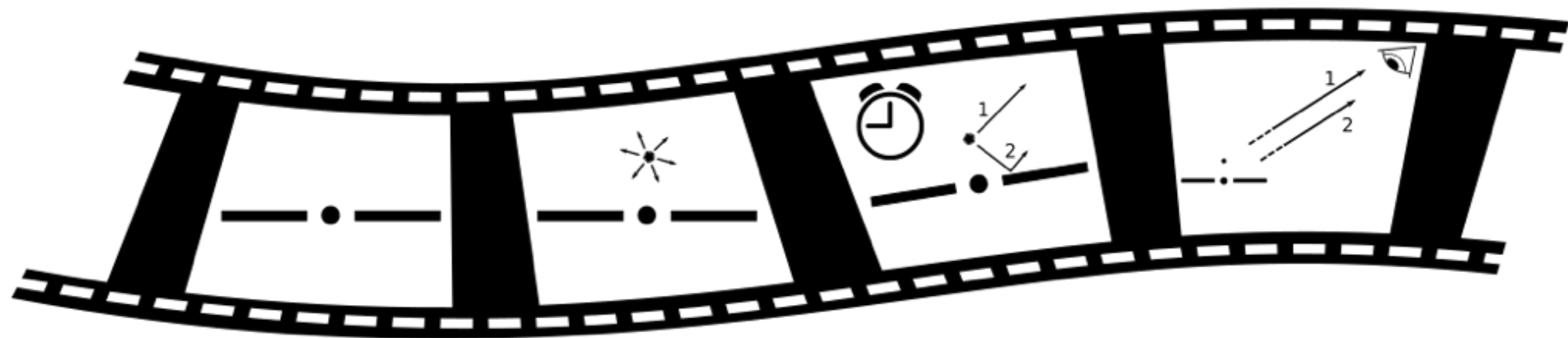




UNIVERSITY OF AMSTERDAM

The effects of spectral hardness changes on reverberation lags

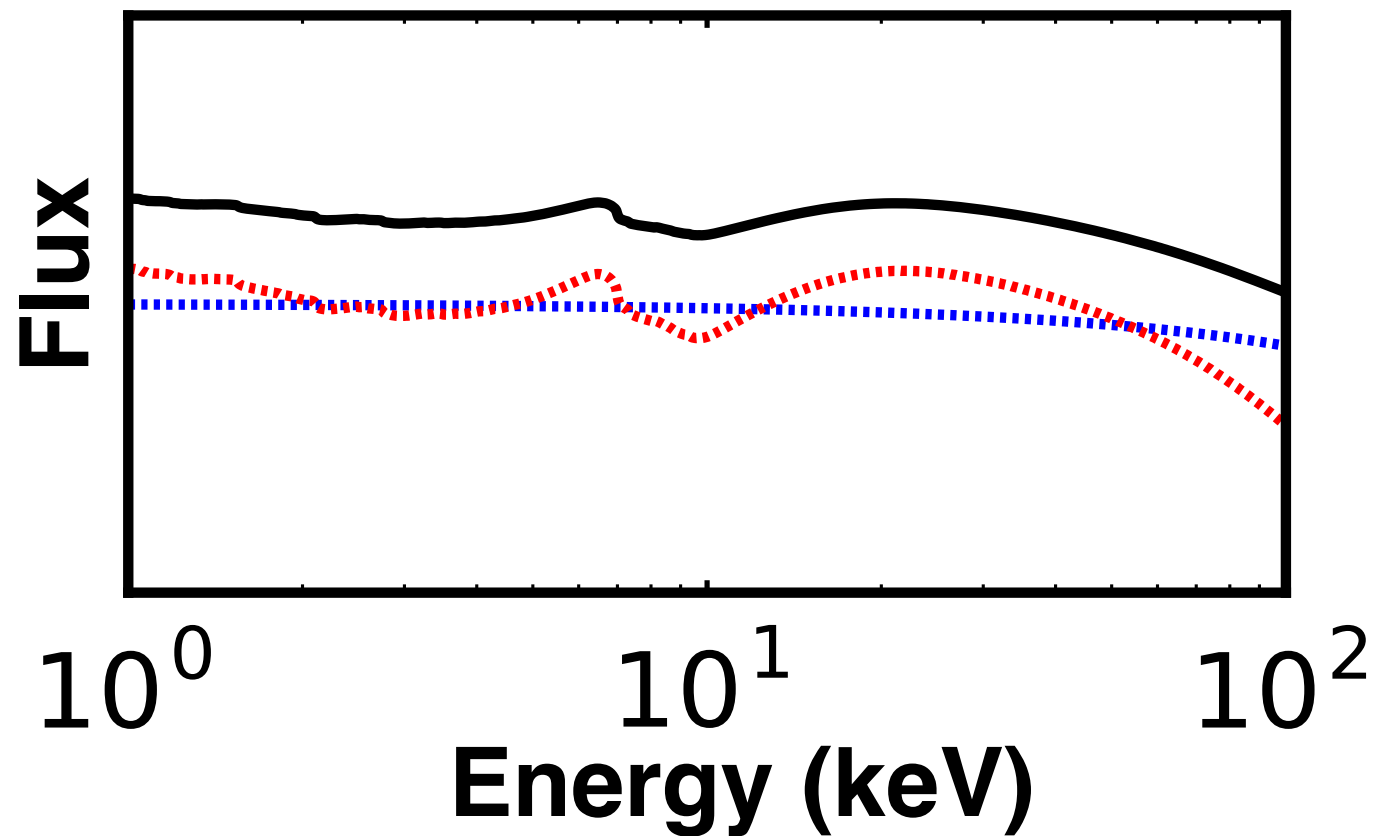
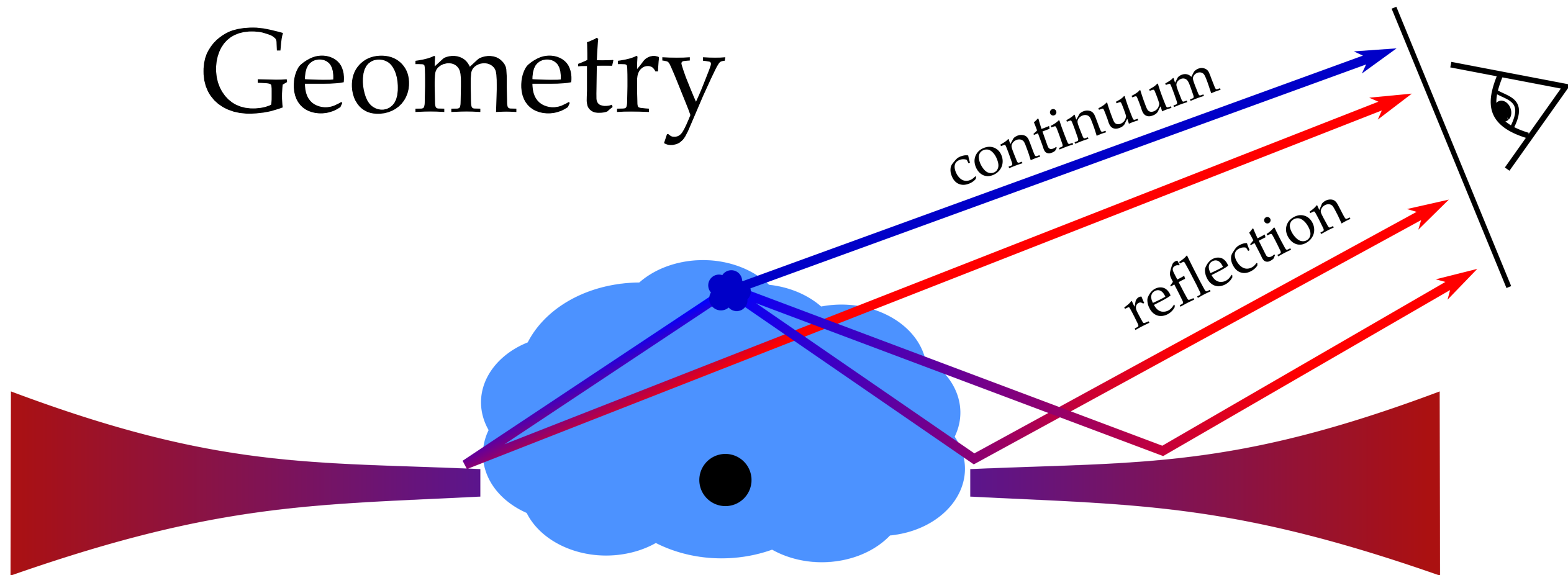
G. Mastroserio, A. Ingram, M. van der Klis



Overview

- Geometry of the systems
- Modelling the lag
- Reverberation with eXTP

Geometry



Inverse Compton scattering

Reflection

Continuum + Reflection

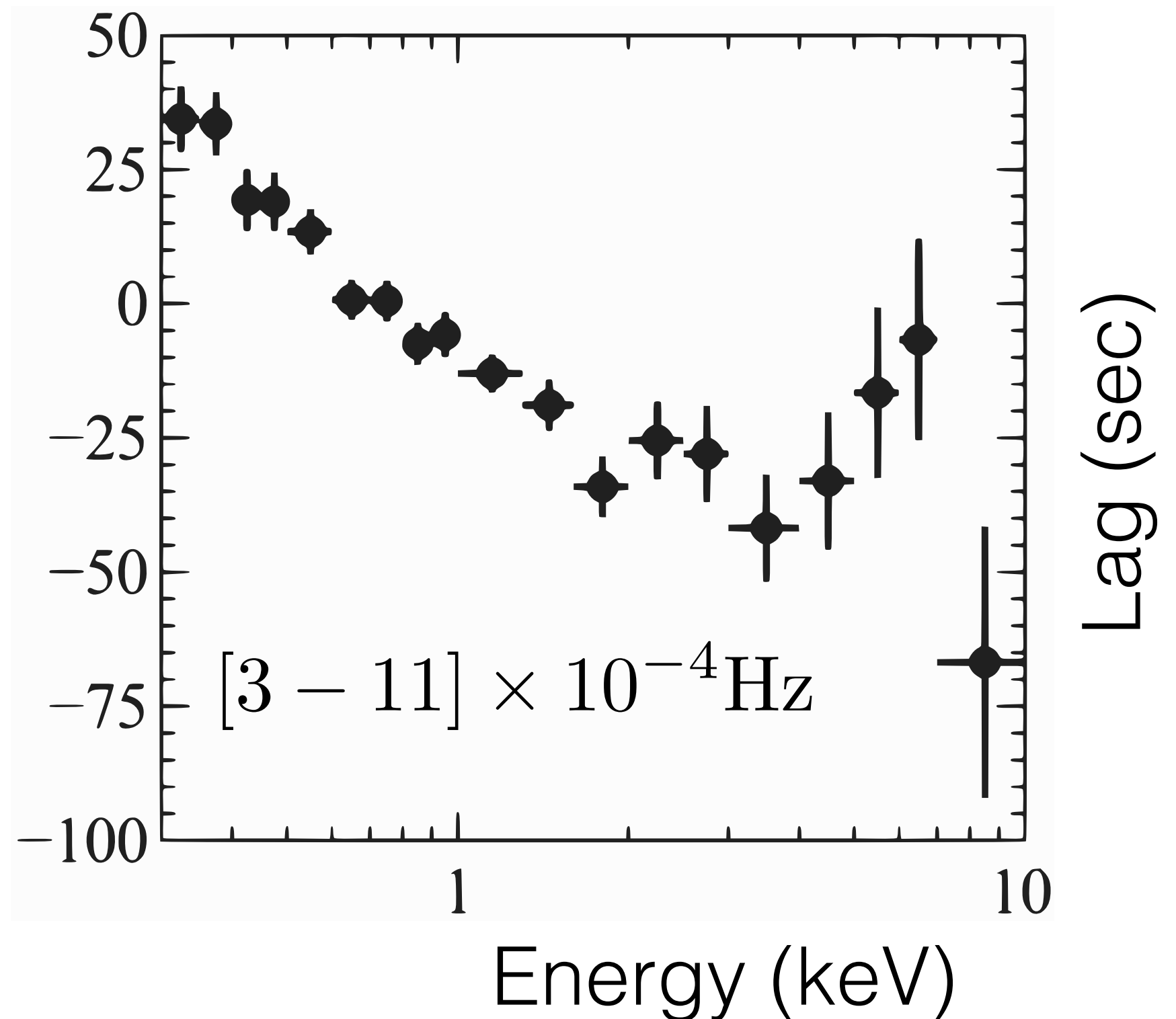
Lag-Energy Spectrum

AGN

Ark 564

$$M \simeq 10^6 M_{\odot}$$

Kara+ 2016



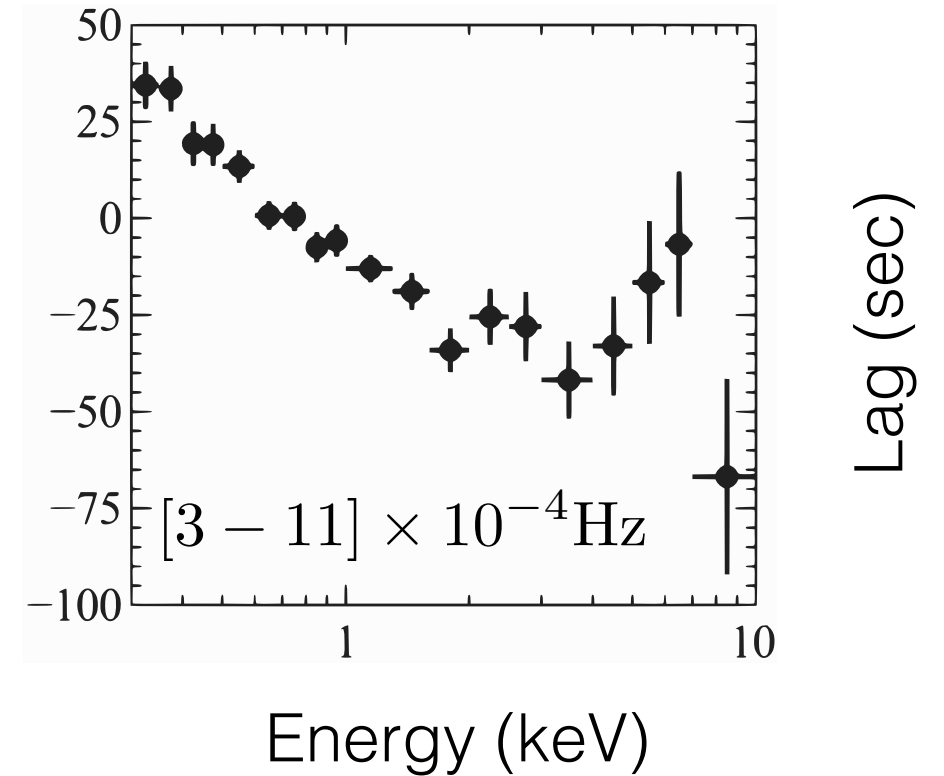
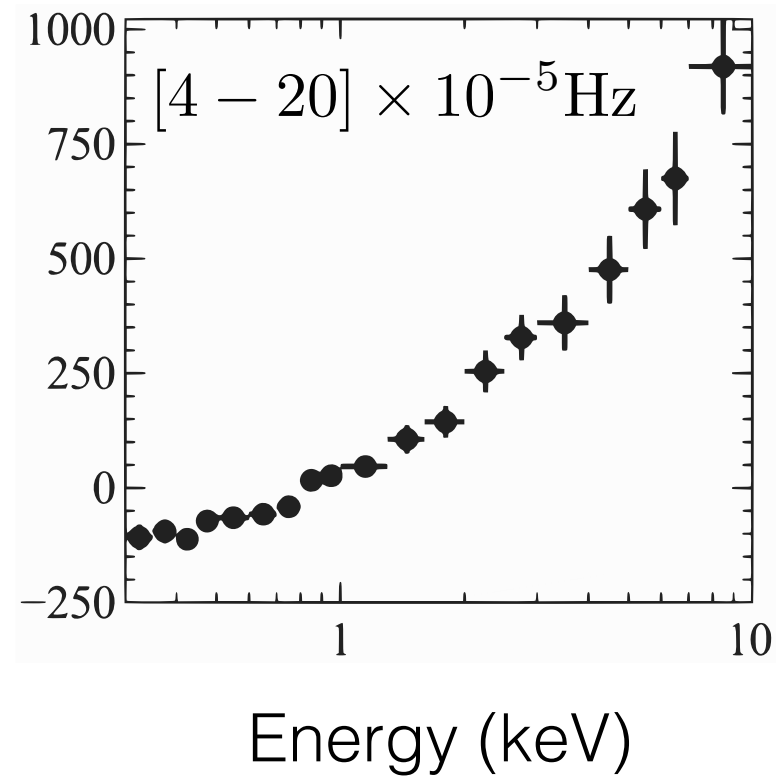
Lag-Energy Spectrum

AGN

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Kara+ 2016



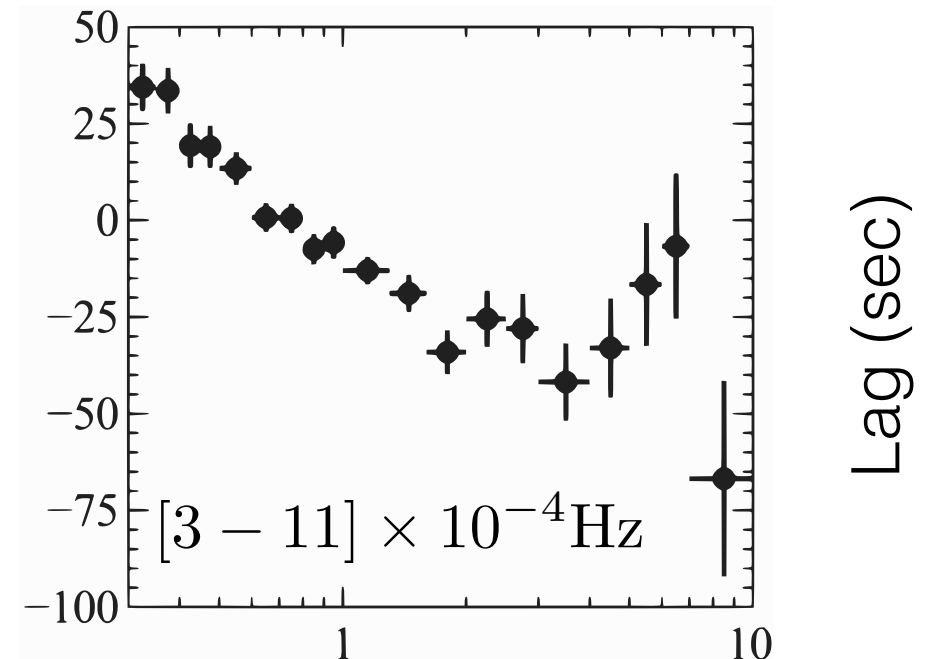
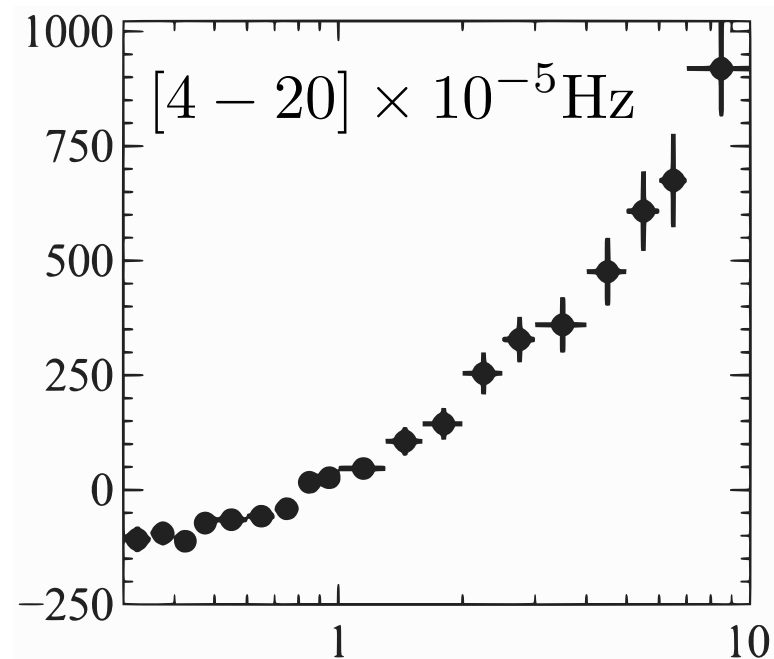
Lag-Energy Spectrum

AGN

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Kara+ 2016

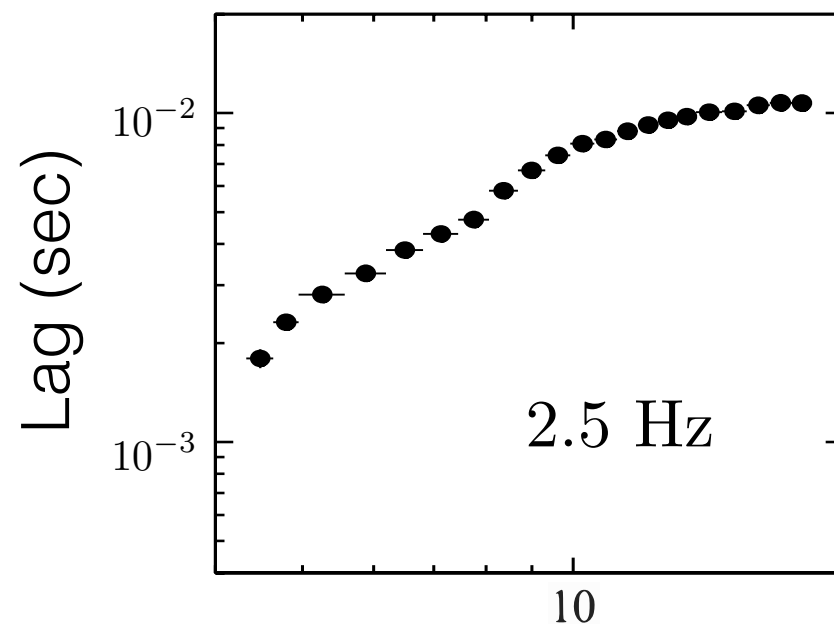


BH binaries

Cyg X-1

$$M \simeq 10 M_{\odot}$$

Kotov+ 2001



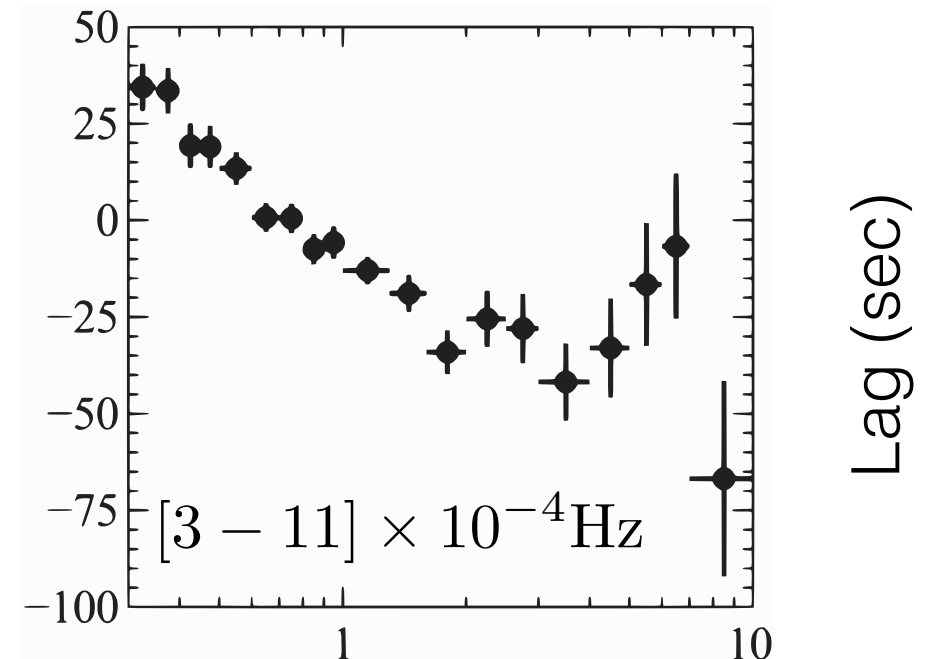
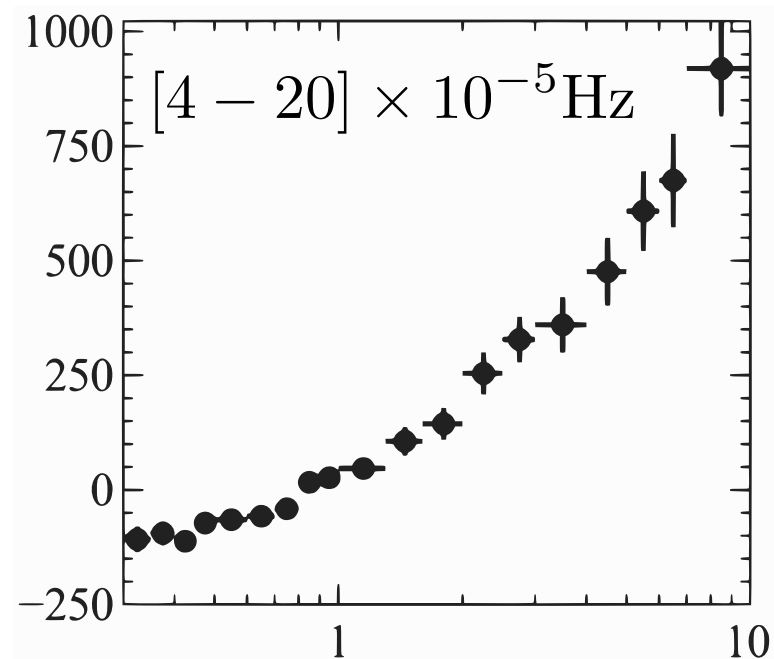
Lag-Energy Spectrum

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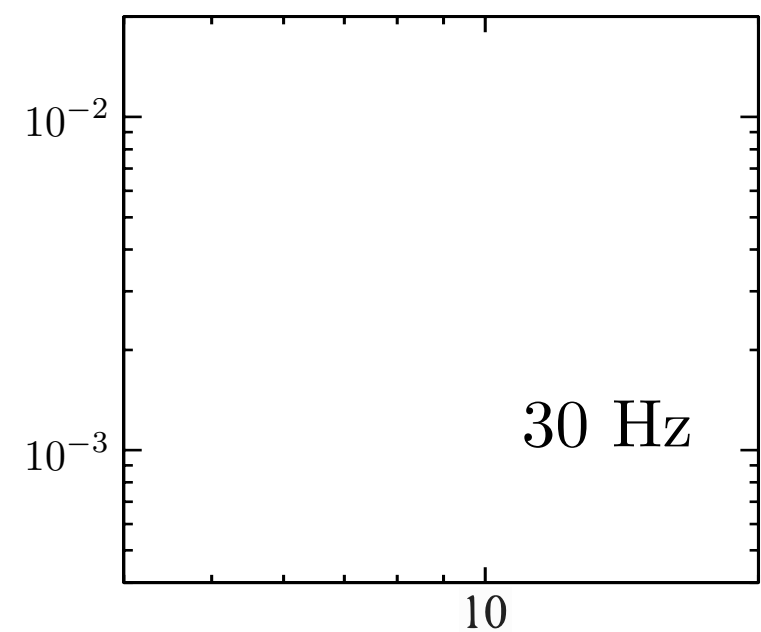
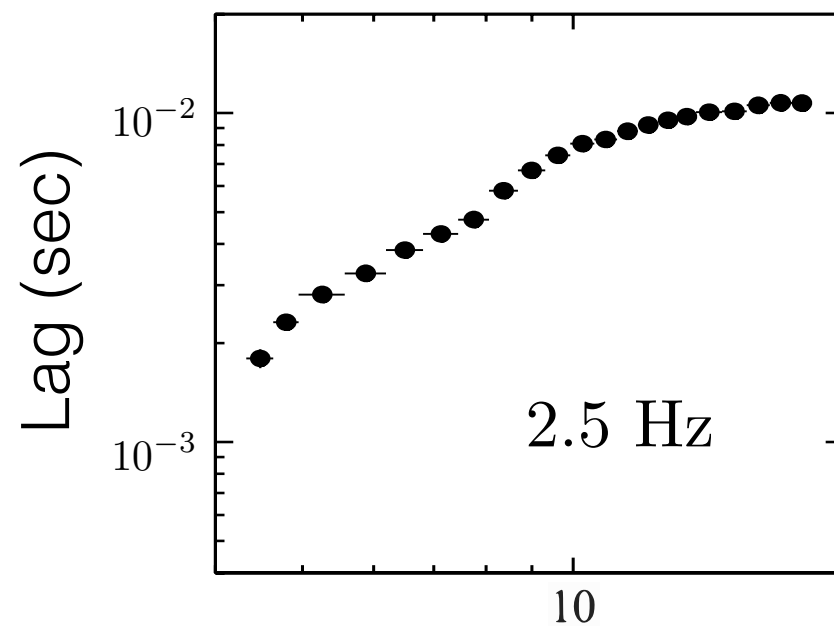


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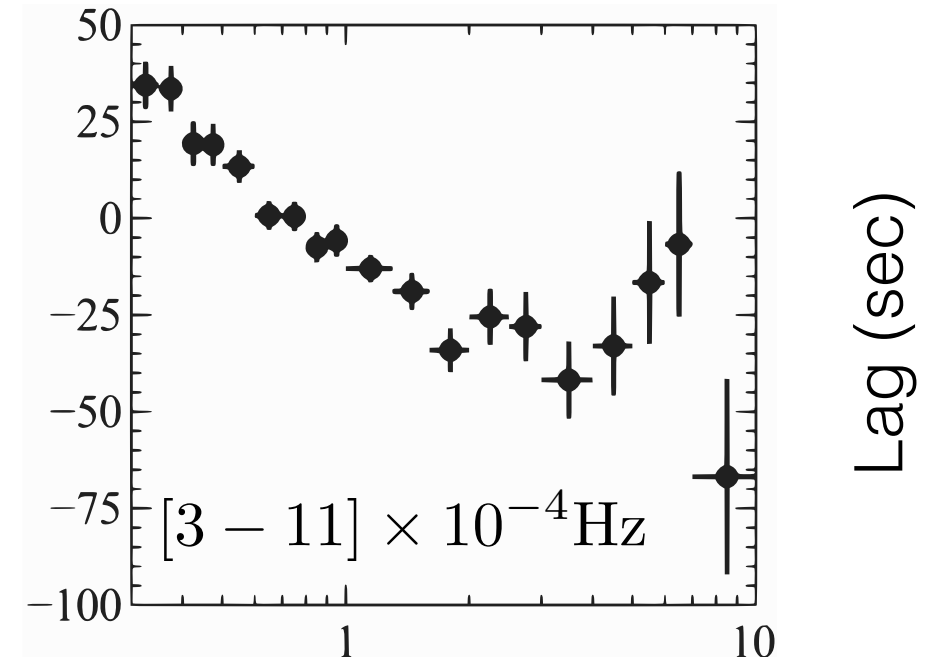
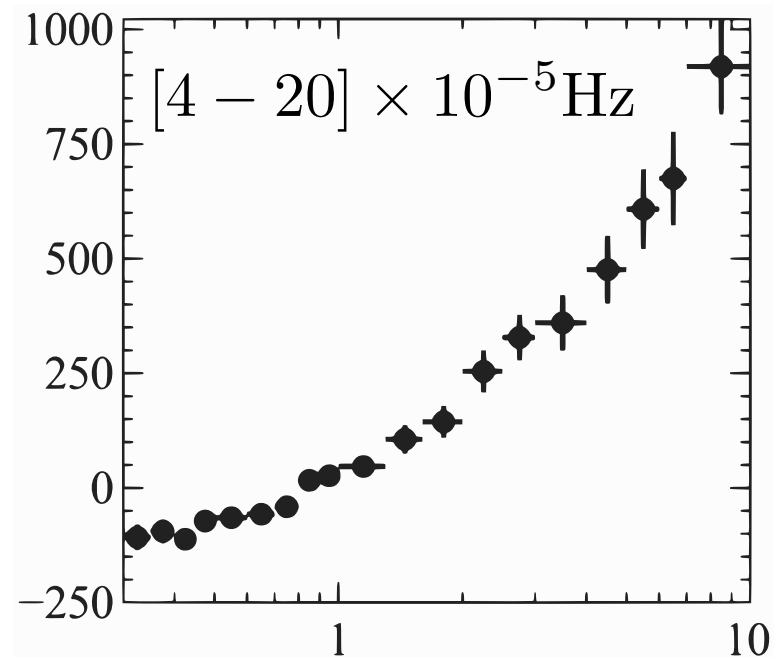
Lag-Energy Spectrum

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Kara+ 2016

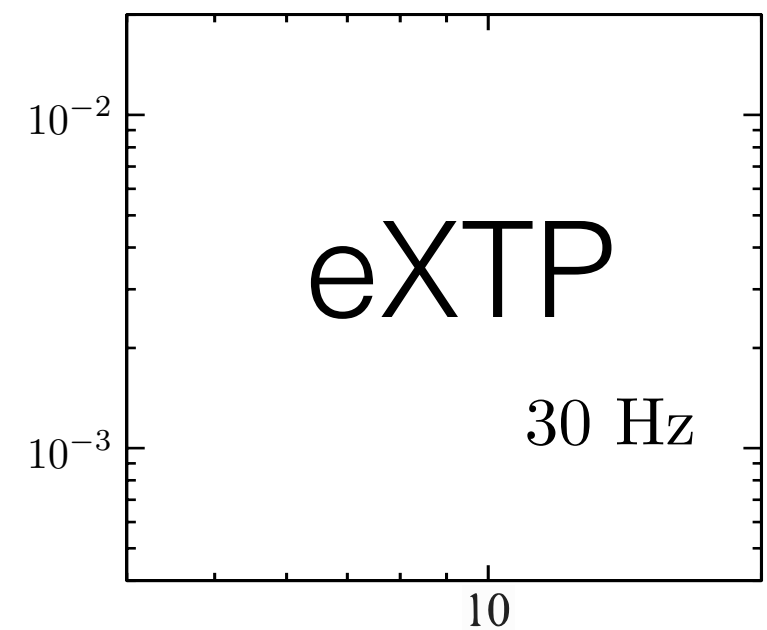
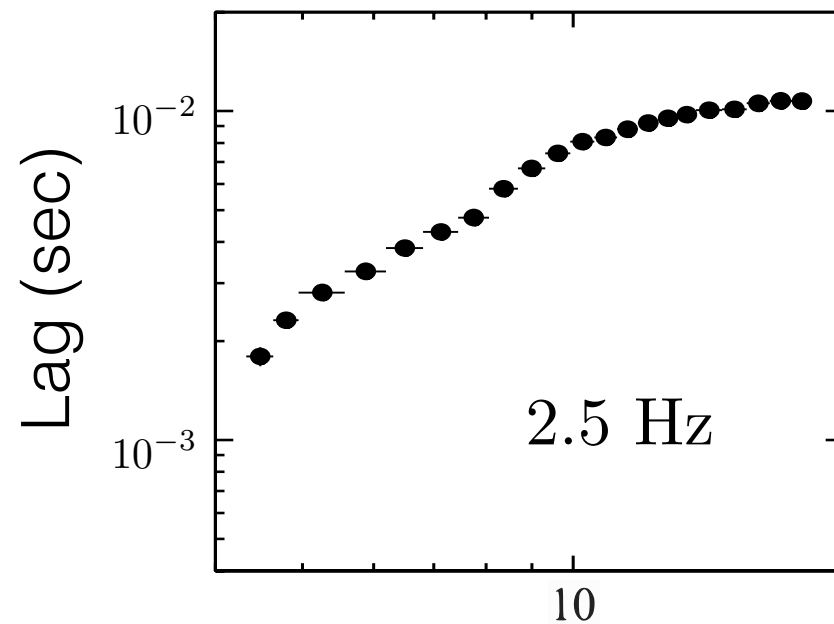


BH binaries

Cyg X-1

$$M \simeq 10 M_{\odot}$$

Kotov+ 2001



Response Function

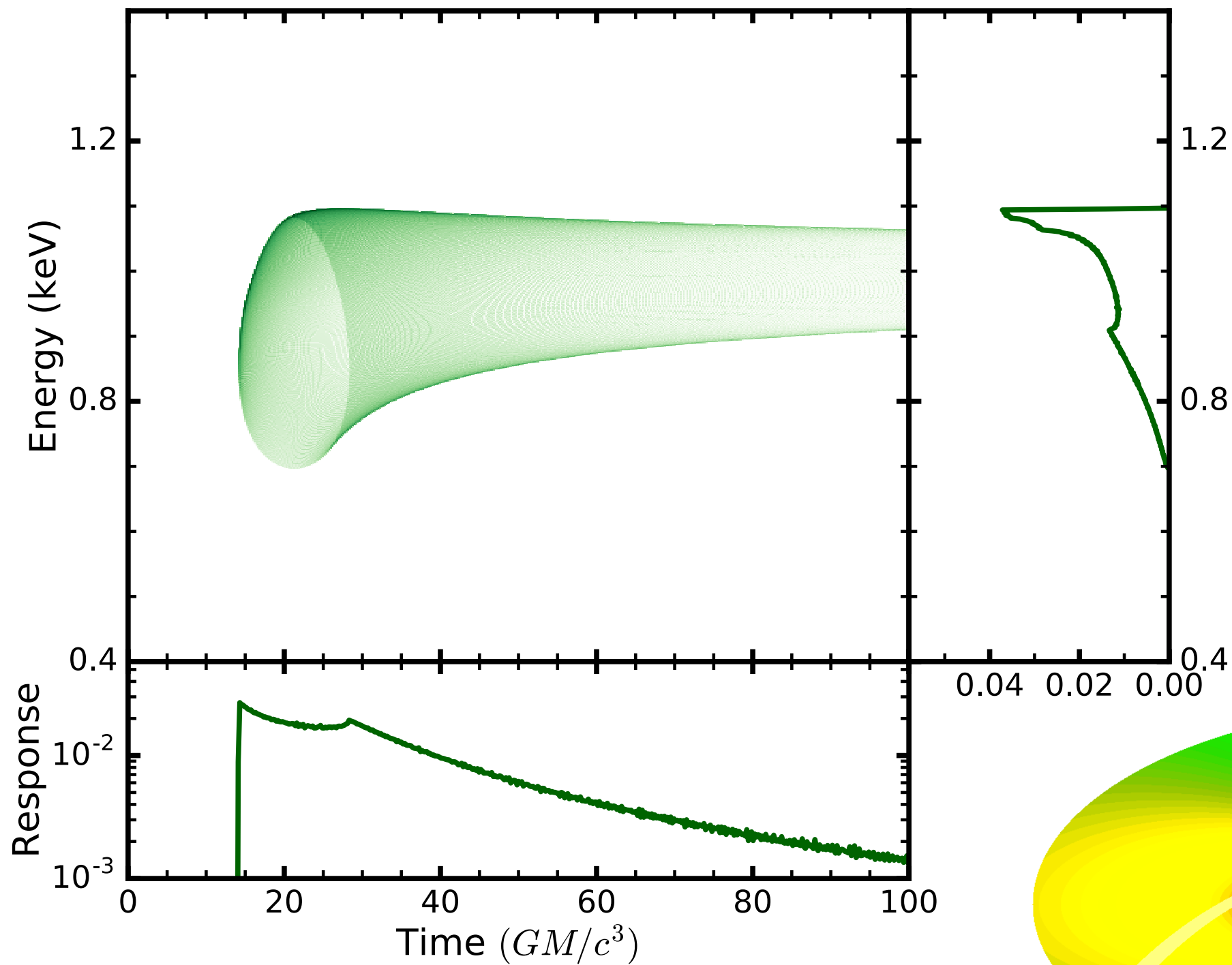
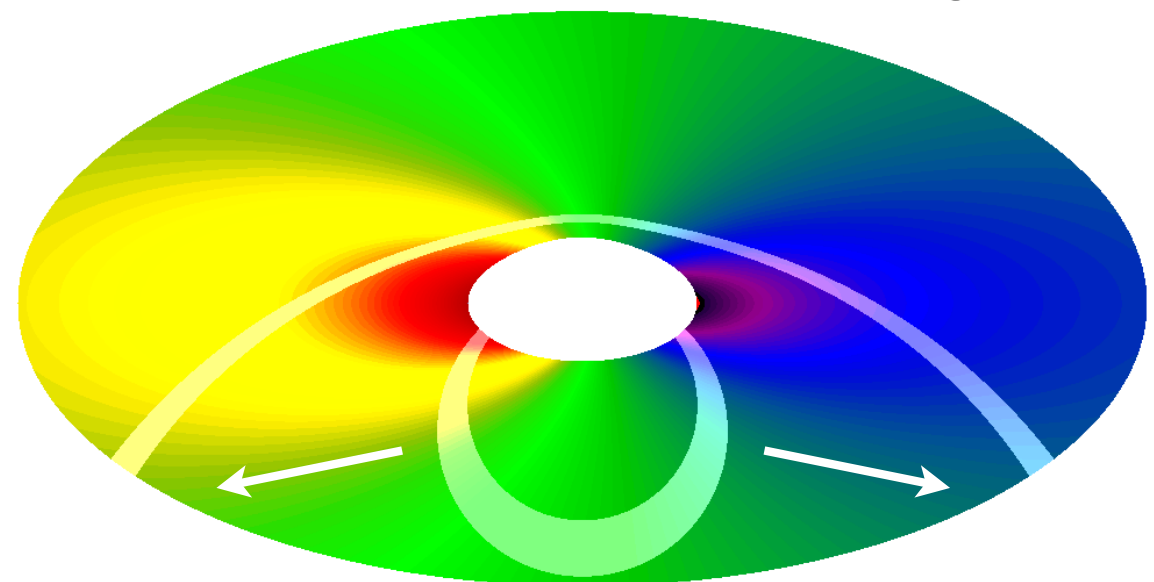
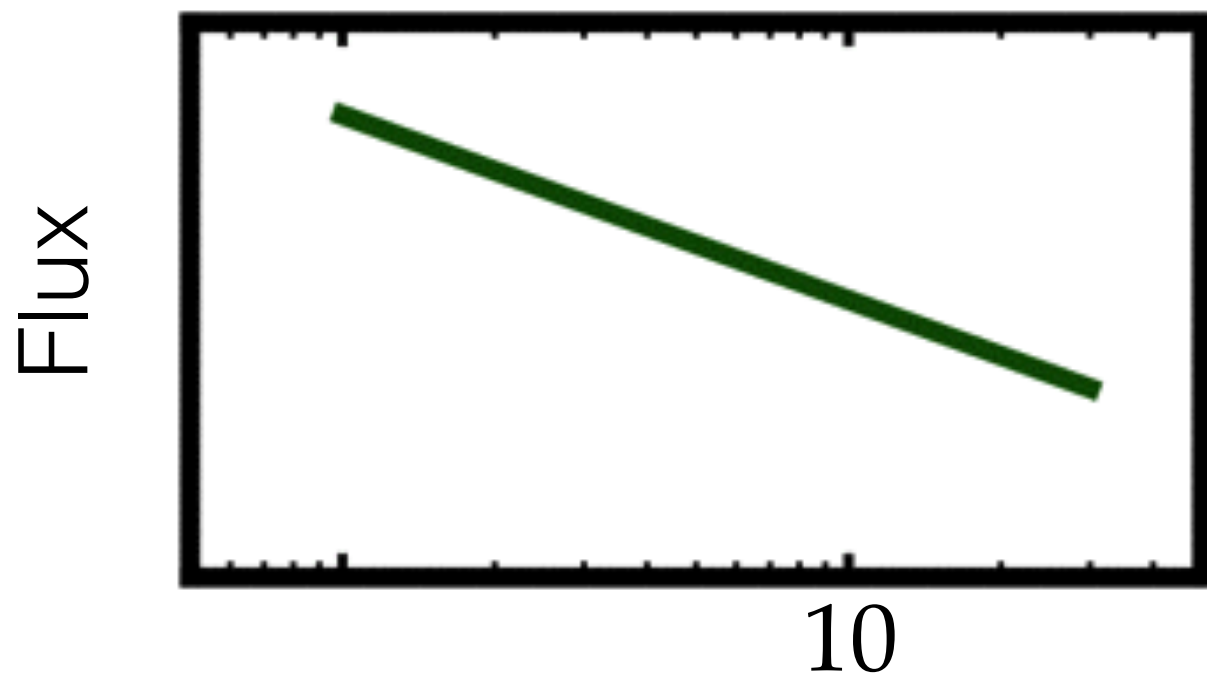


Image credit:
A. Ingram

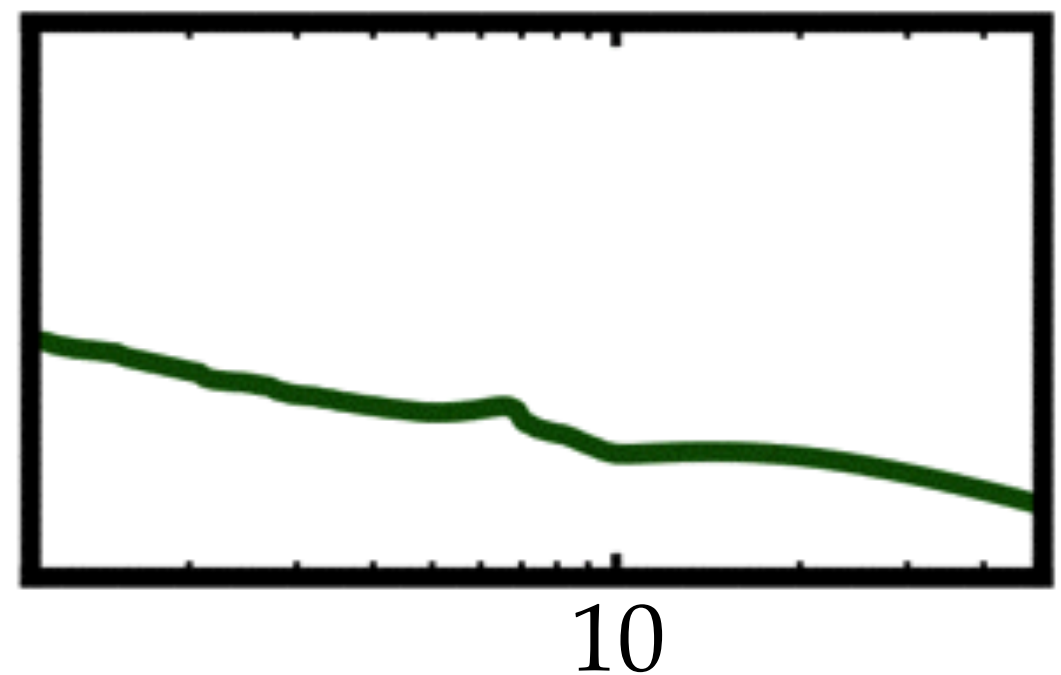


Previous models

Continuum



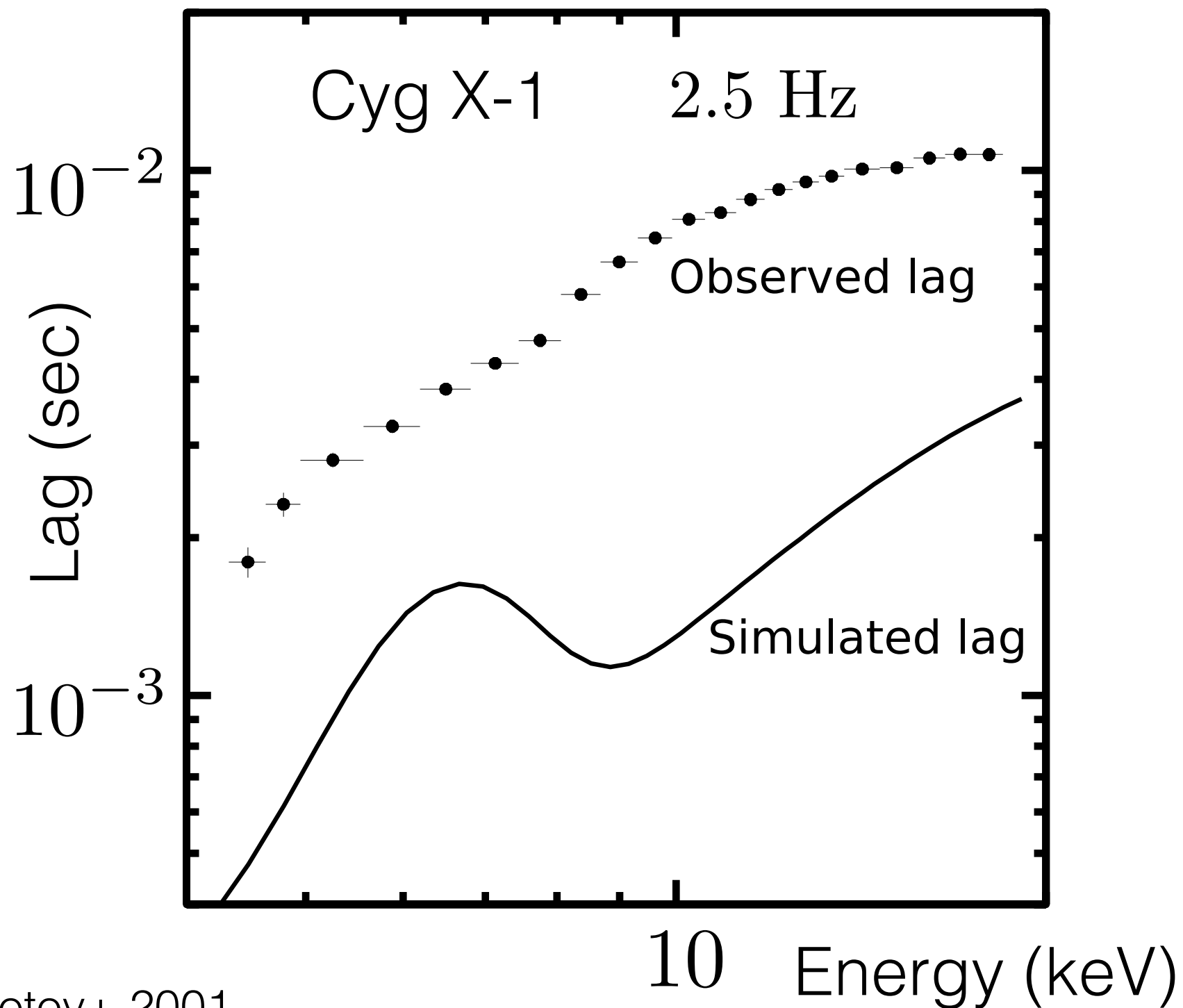
Reflection



Energy (keV)

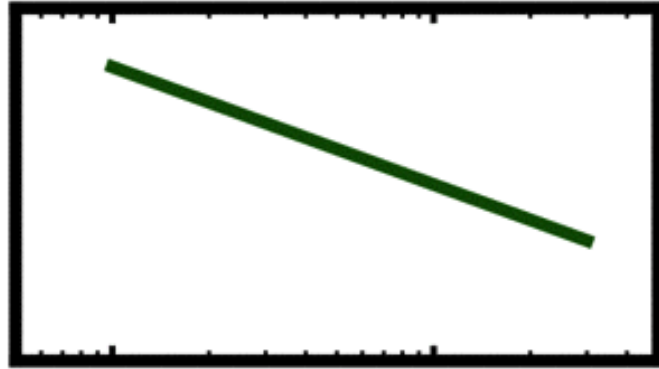
Previous models

for black hole binaries

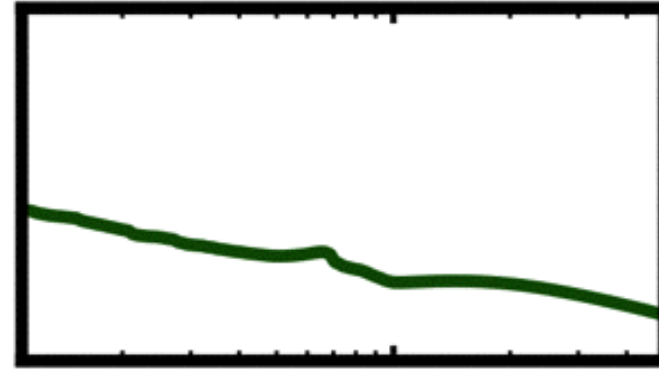


Kotov+ 2001

Spectral Hardness Changes

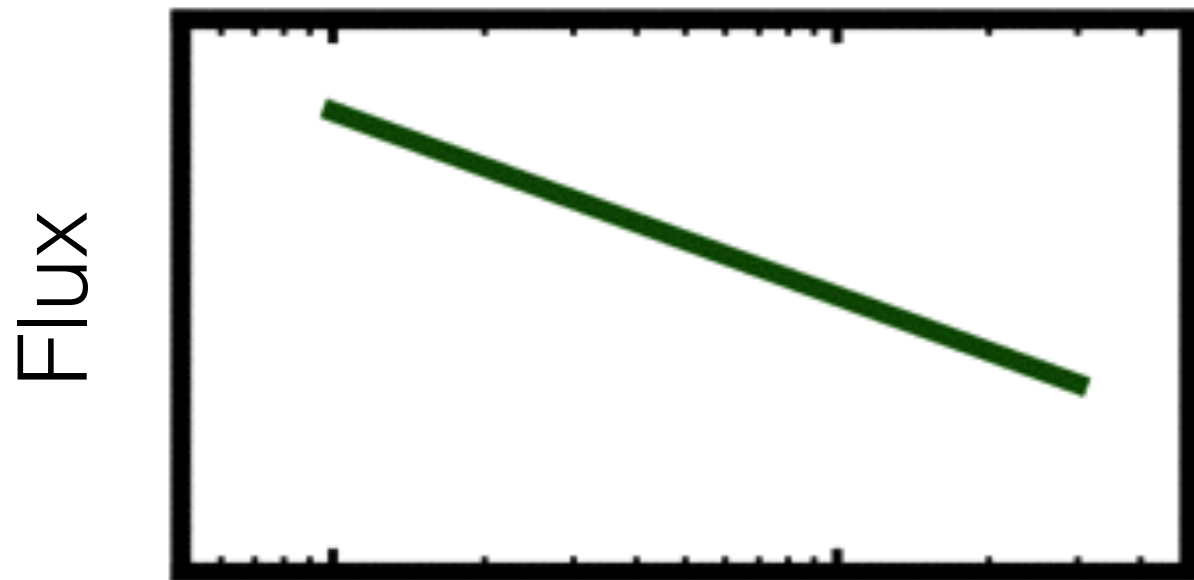


Continuum



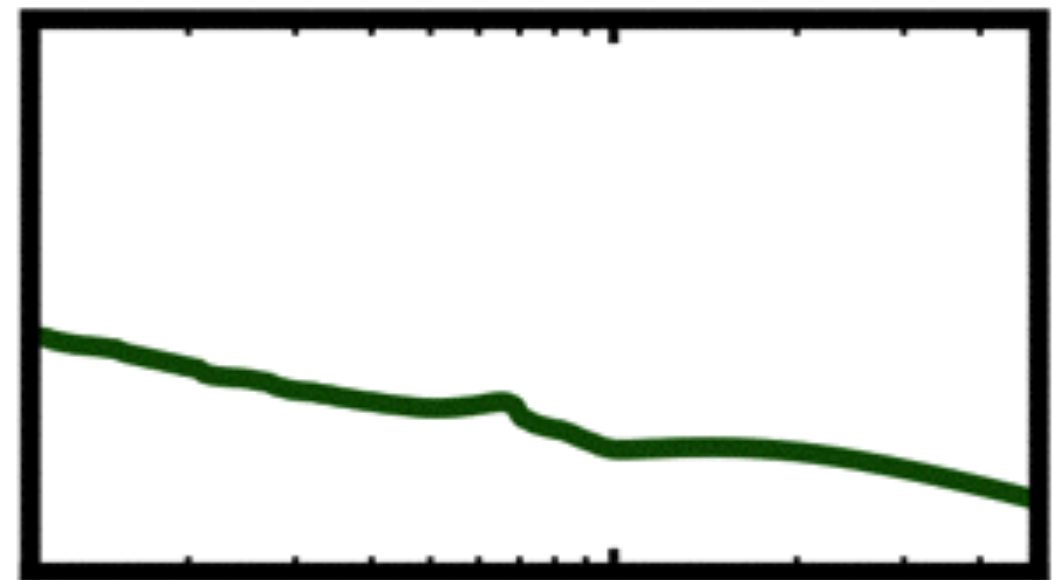
Reflection

PIVOTING



10

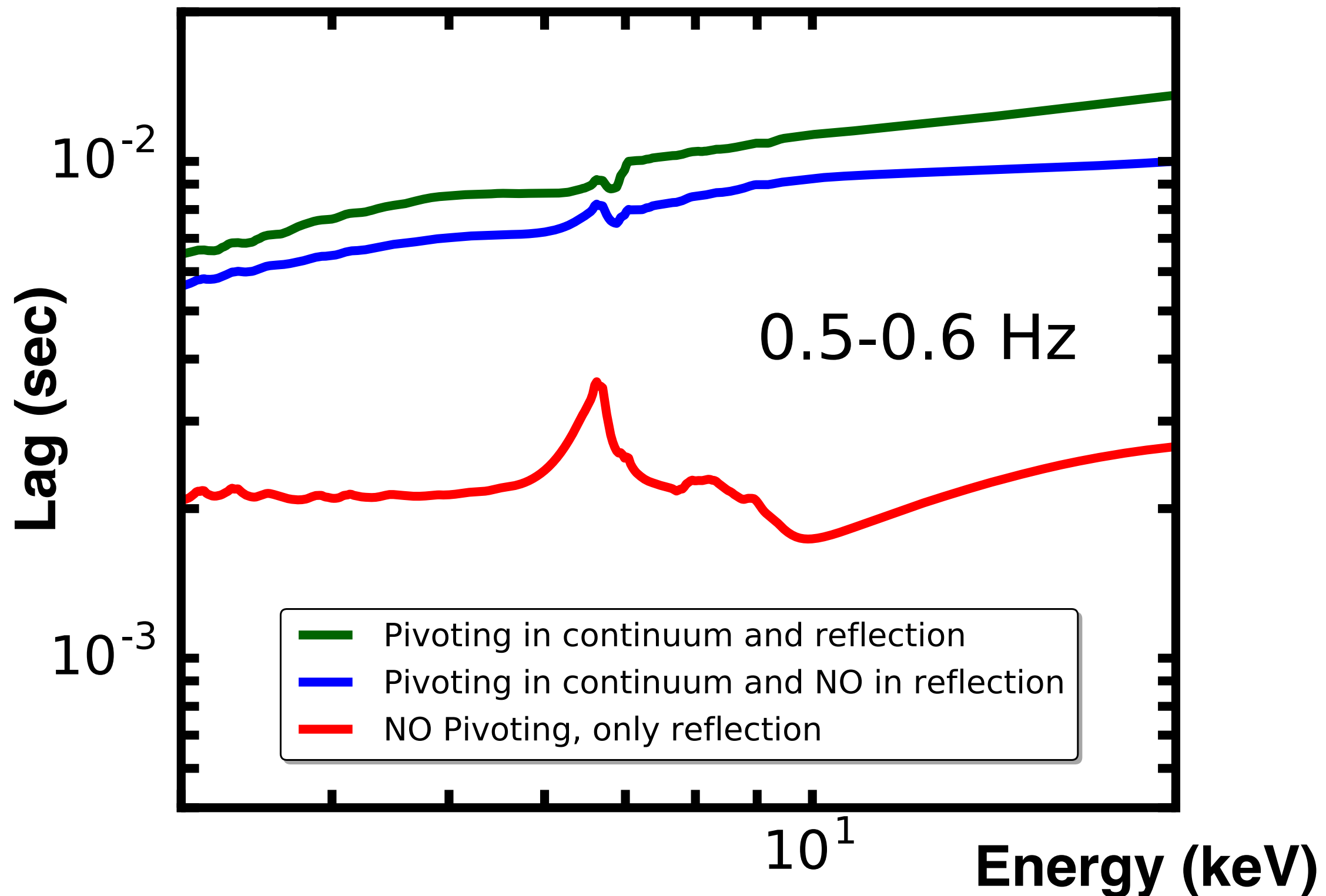
Energy (keV)



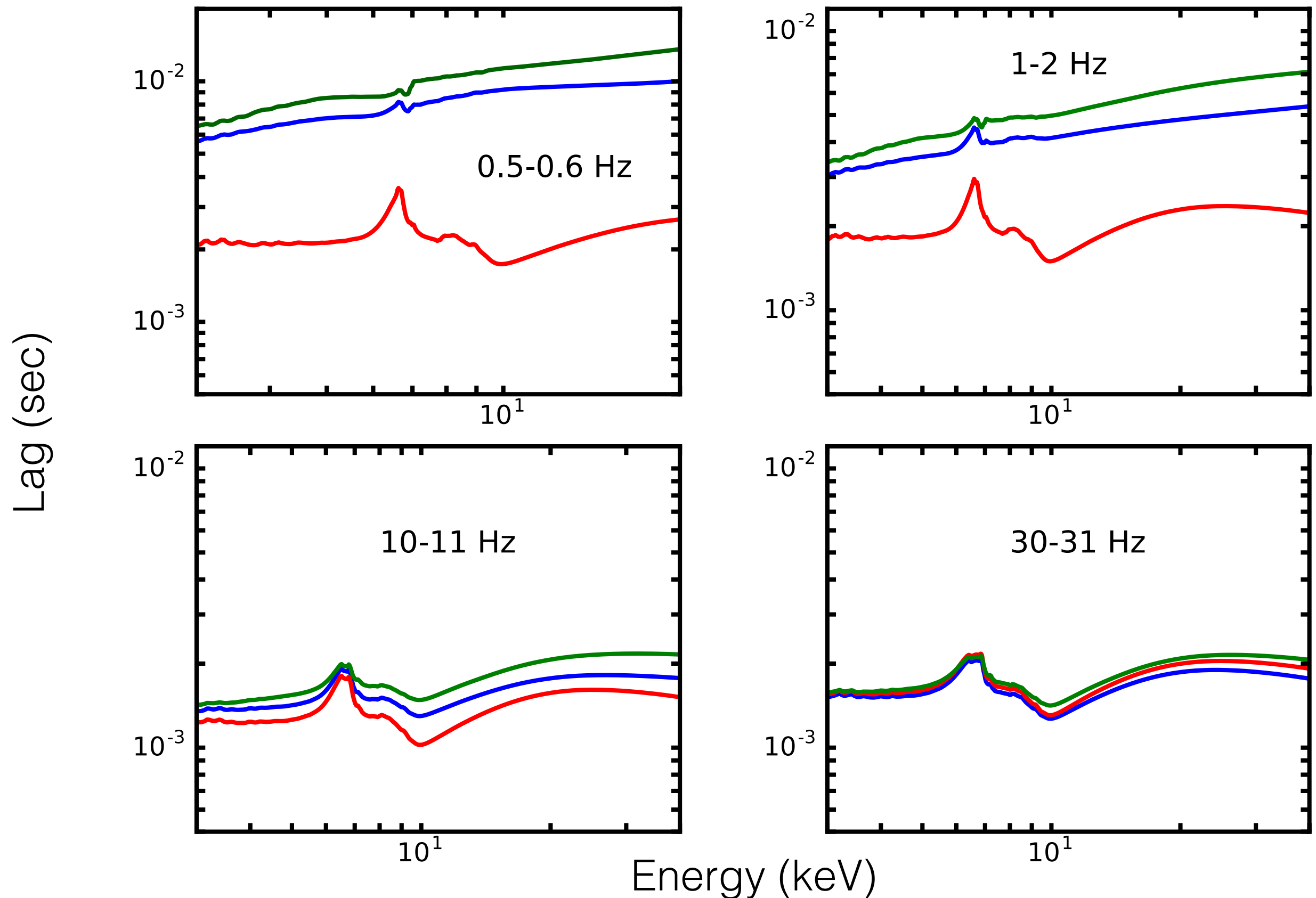
10

Pivoting Model

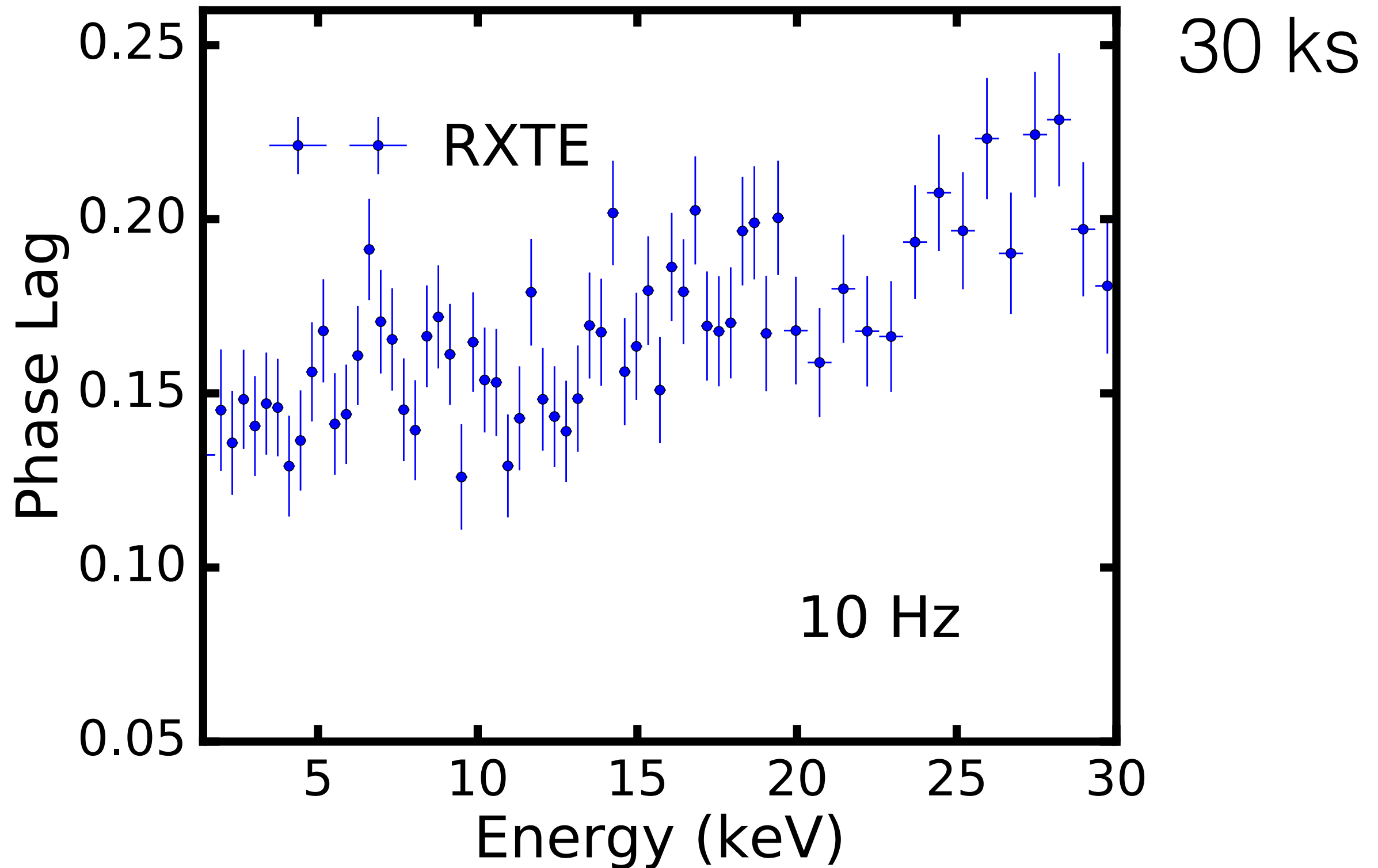
for black hole binaries



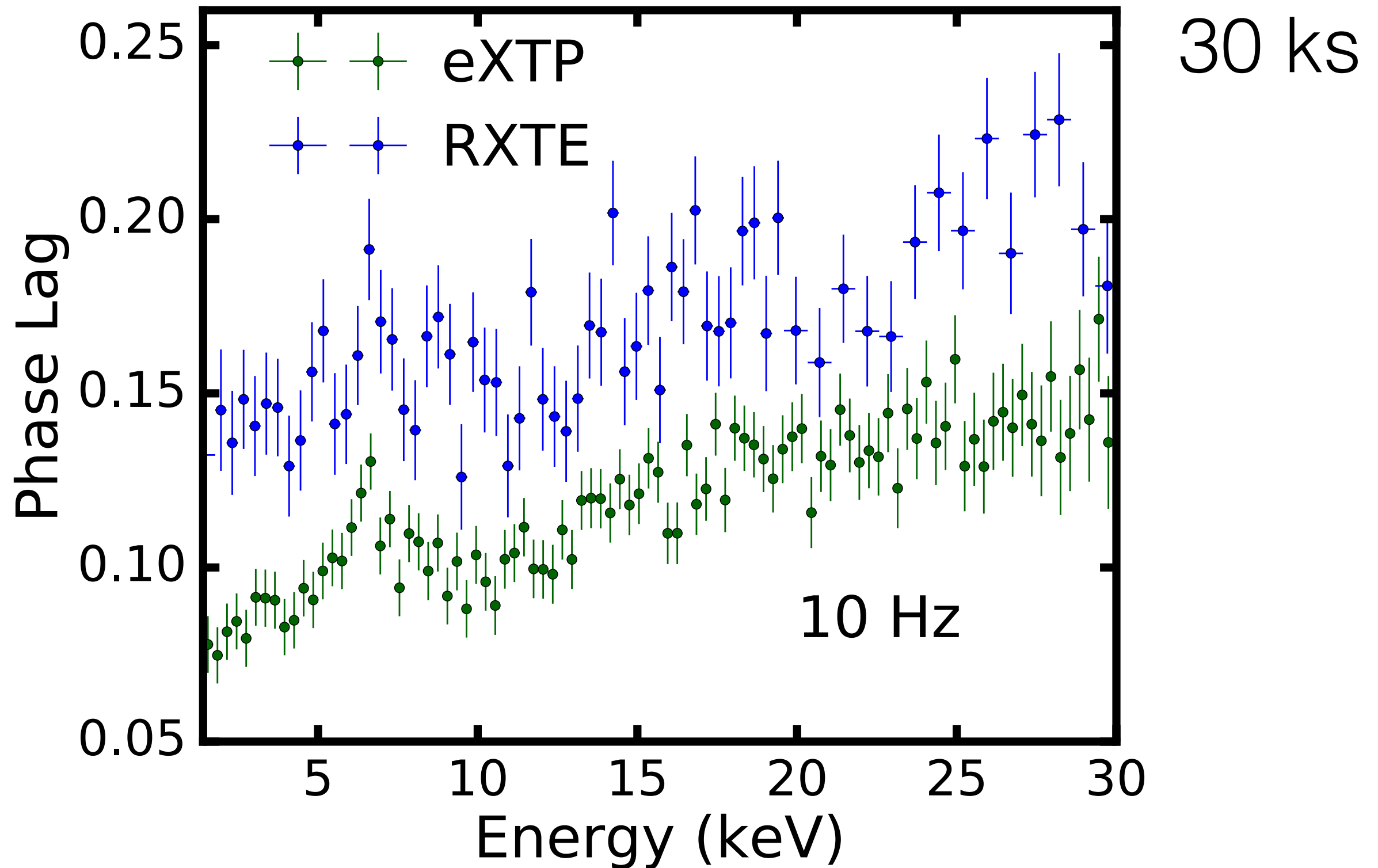
Pivoting Model



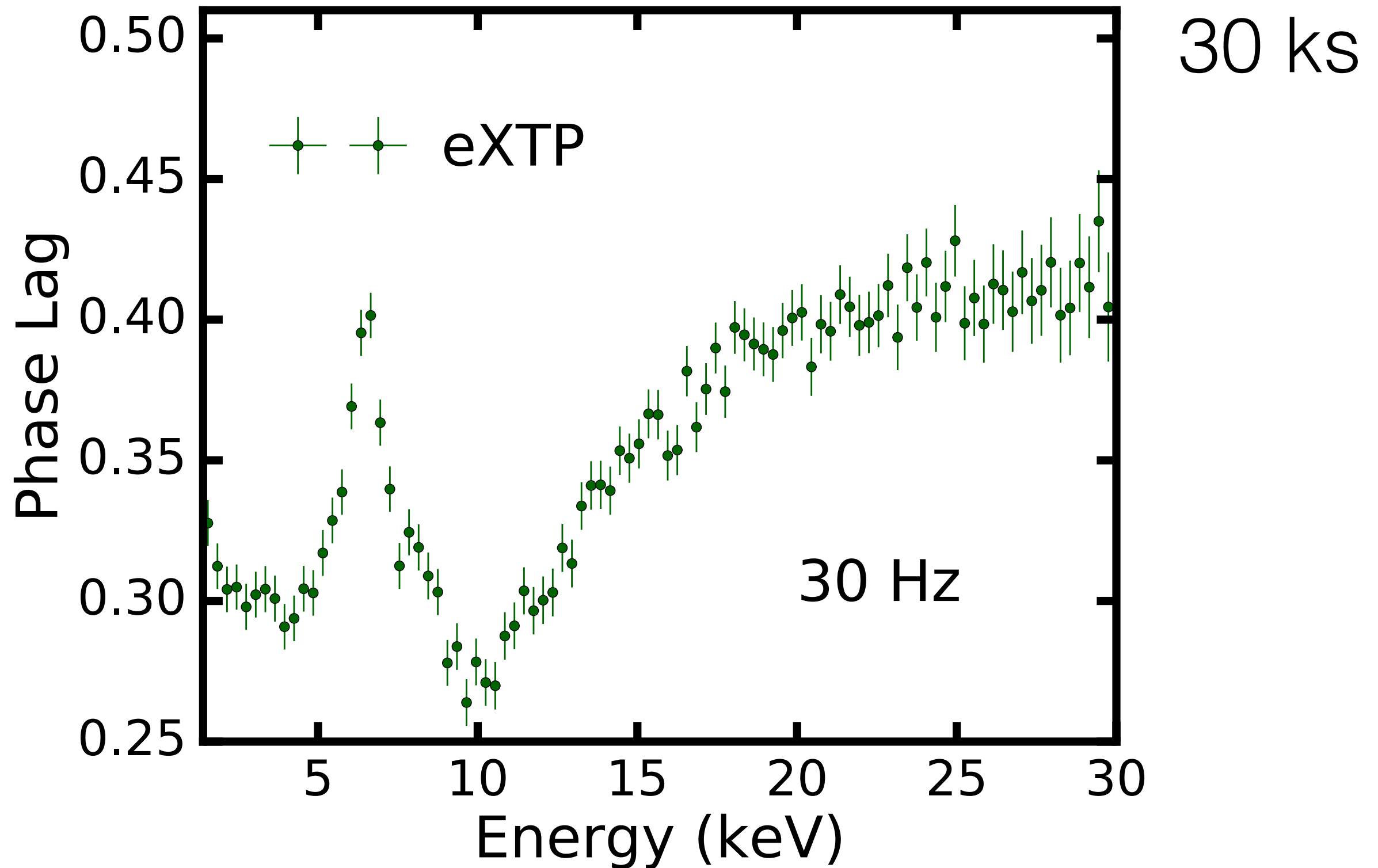
eXTP Simulations



eXTP Simulations



eXTP Simulations



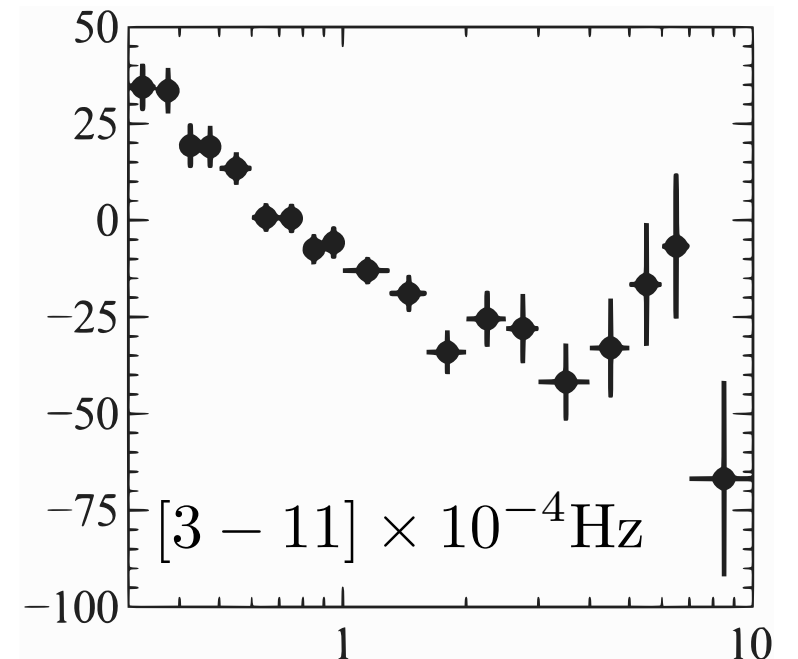
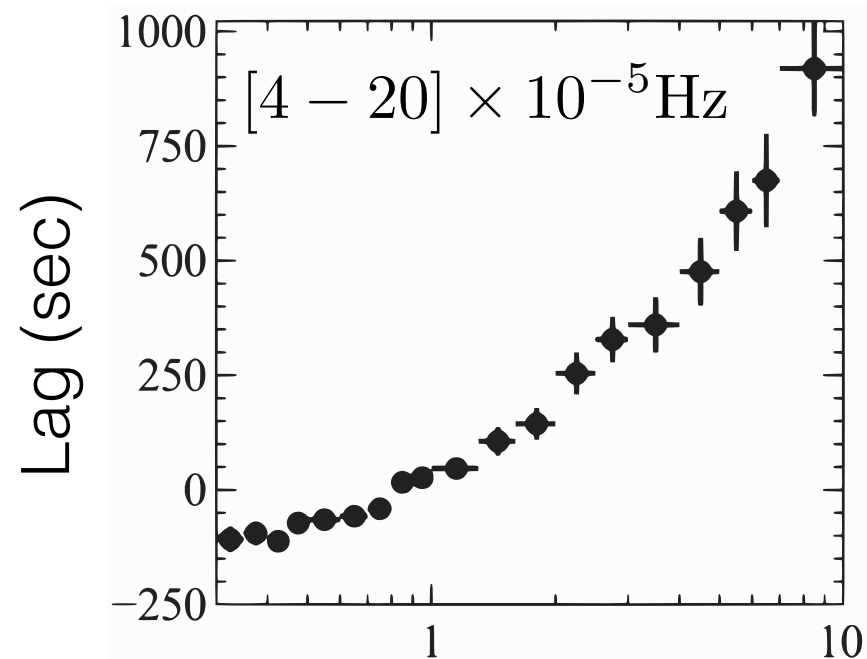
Lag-Energy Spectrum

AGN

Ark 564

$$M \simeq 10^6 M_{\odot}$$

Kara+ 2016

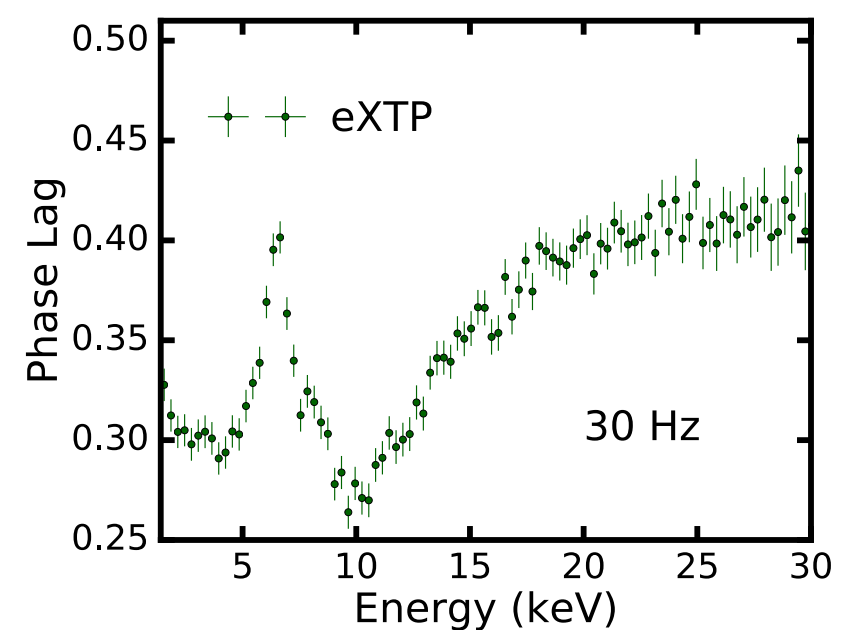
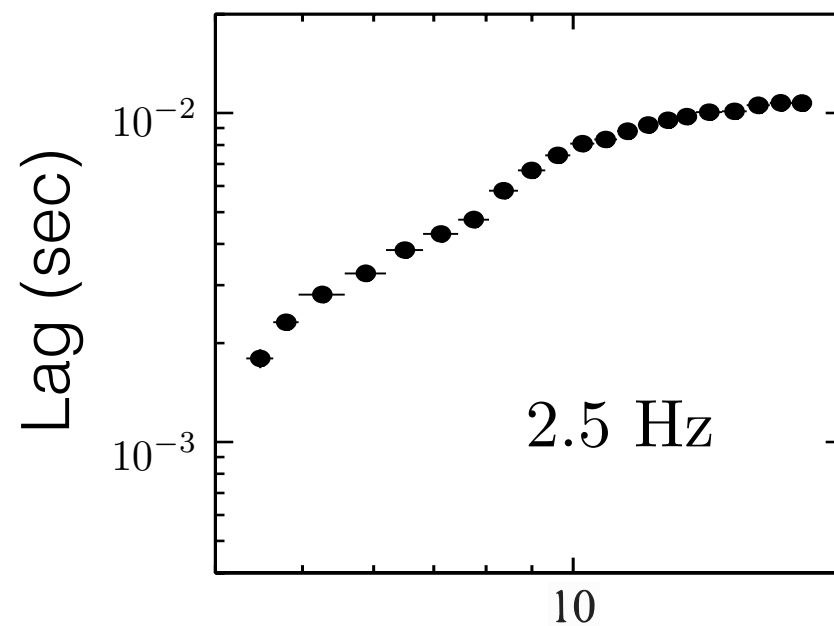


BH binaries

Cyg X-1

$$M \simeq 10 M_{\odot}$$

Kotov+ 2001

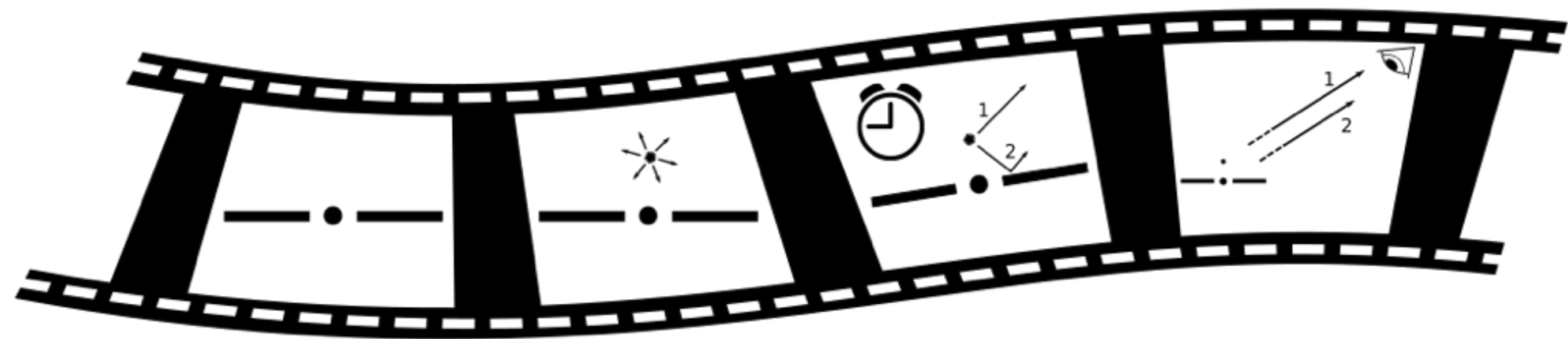


Conclusions

The new model explains both the continuum and the reflection lag consistently

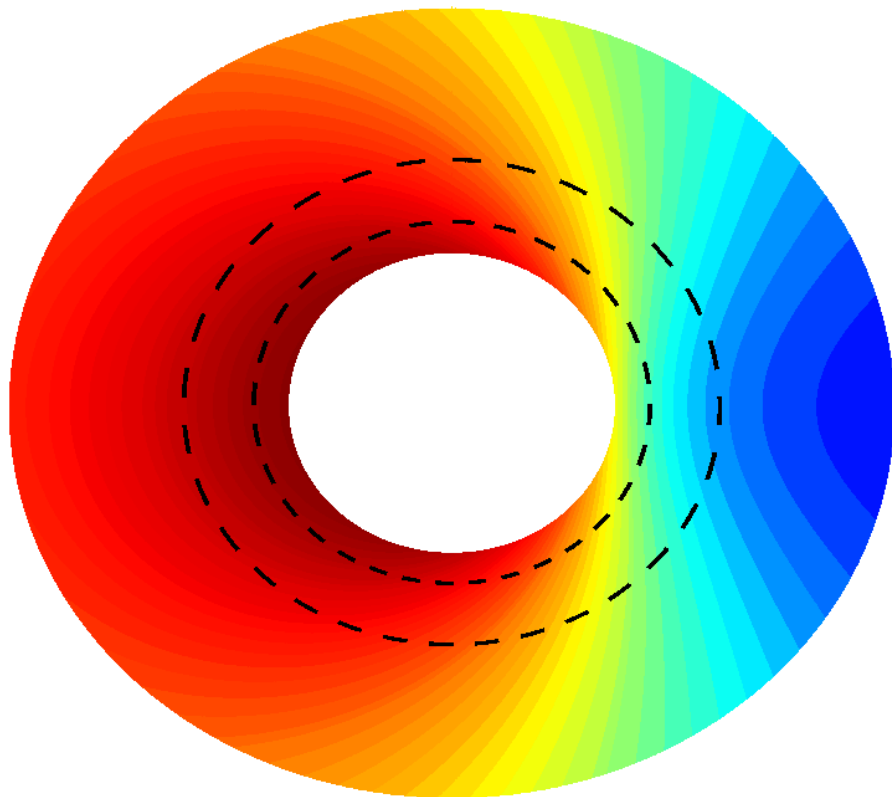
- The model depends on the disk geometry and black hole parameters
- The model is analytical and very flexible for introducing new components
- We can fit our model with the cross-spectrum not only with the lag spectrum
- eXTP will provide a great improvement in the detection of reverberation lag

Thank you



K α Iron Line Profile

Fabian+ 2000

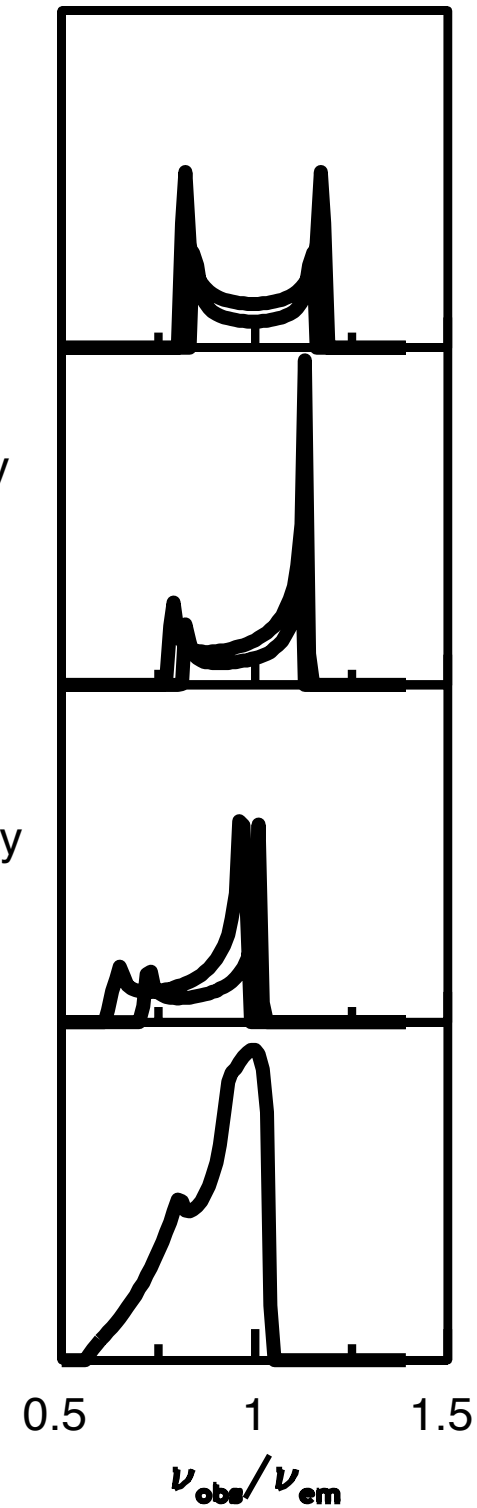


Newtonian

Special relativity

General relativity

Line profile



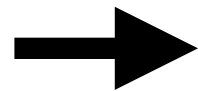
Transverse Doppler shift

Beaming

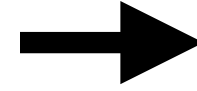
Gravitational redshift

Transfer Function

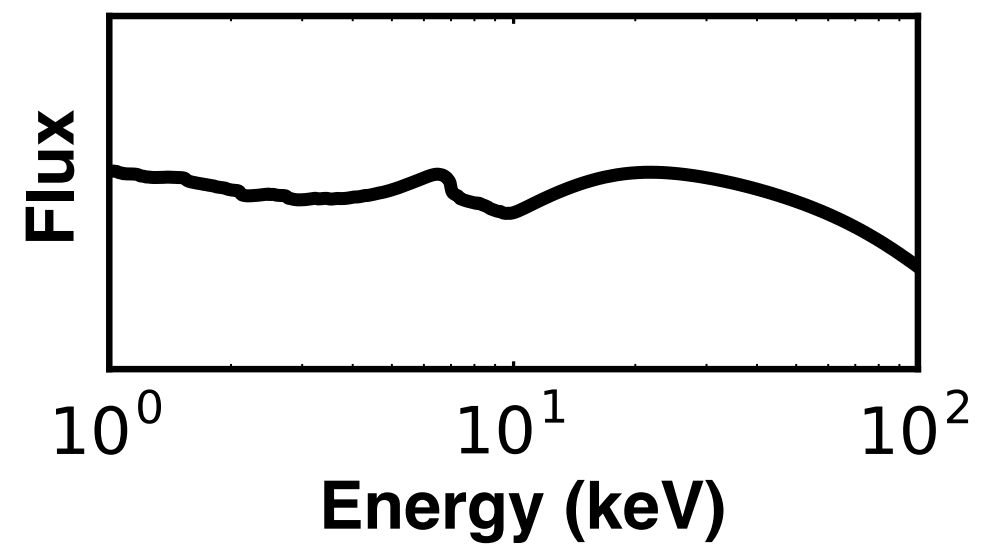
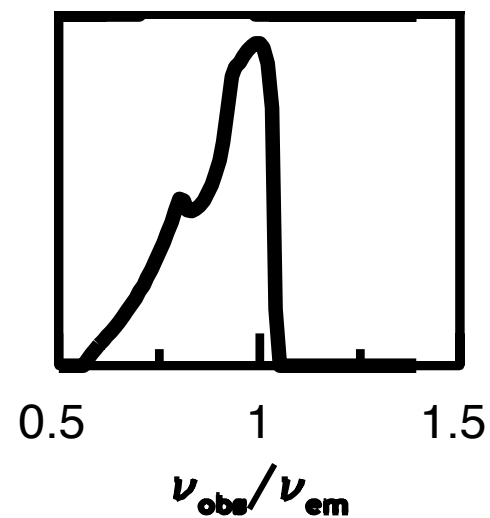
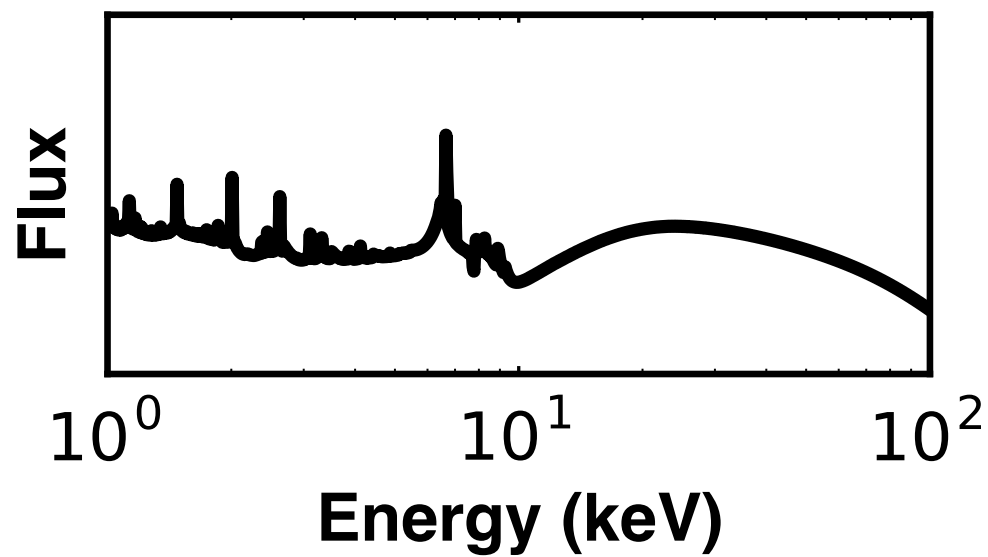
Rest frame
spectrum



Transfer function

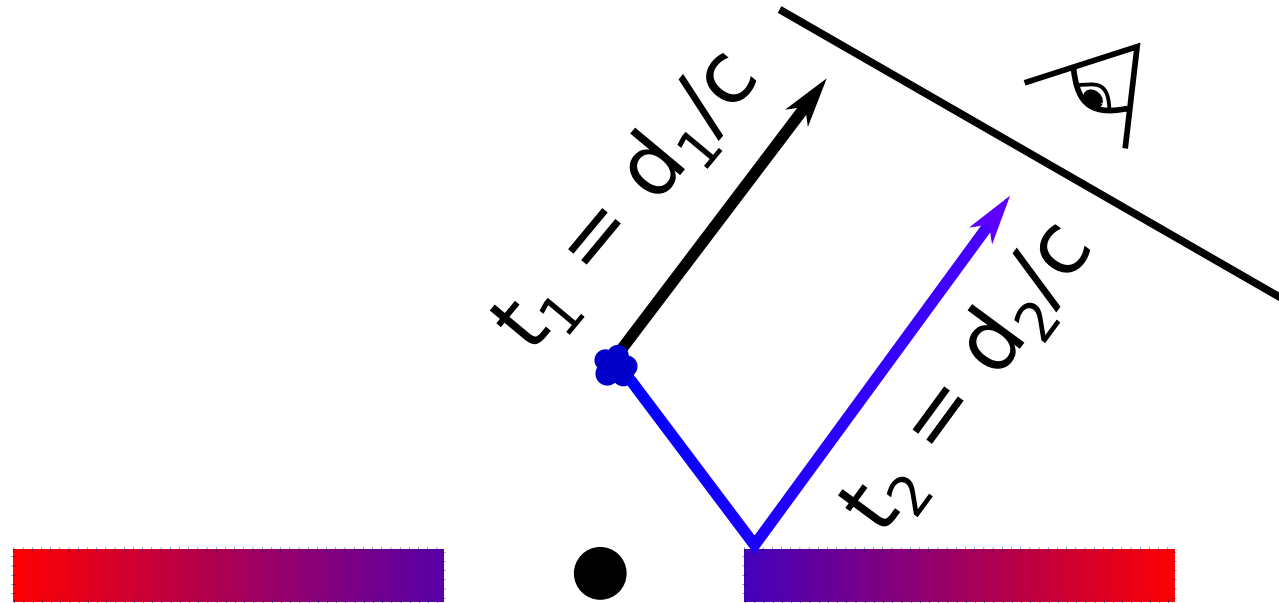


Reflection
spectrum

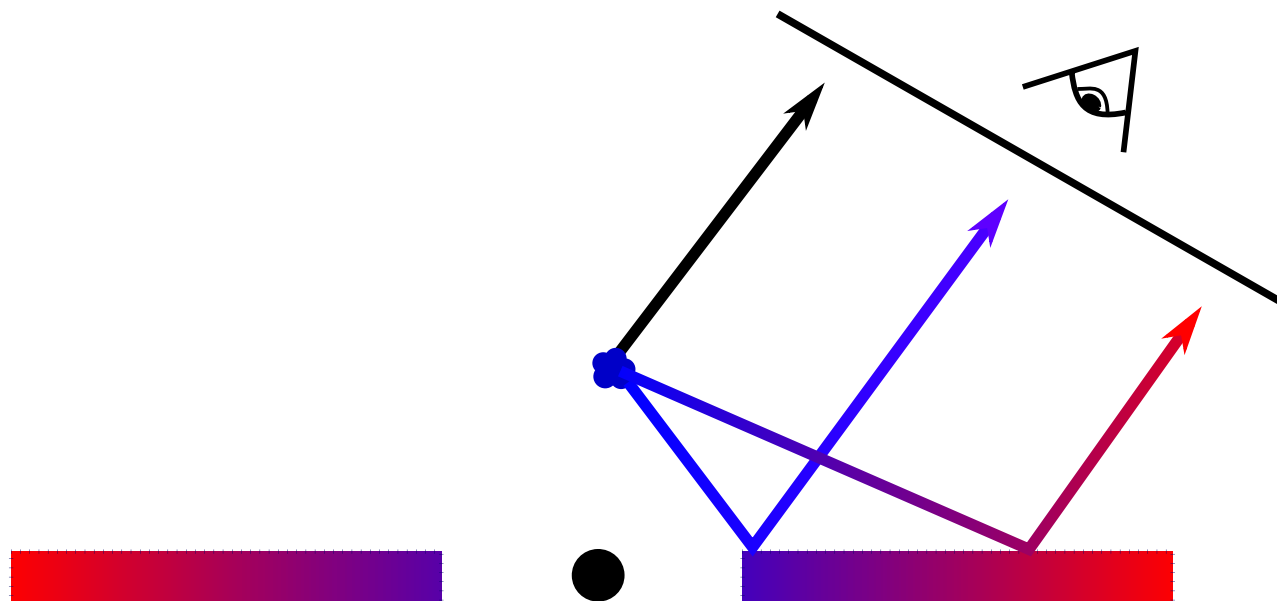


XILLVER model
García&Kallman 2010

Reverberation Lag



$$\Delta t = t_1 - t_2$$



Reverberation lag

