

# A Sub-keV X-ray Polarimeter

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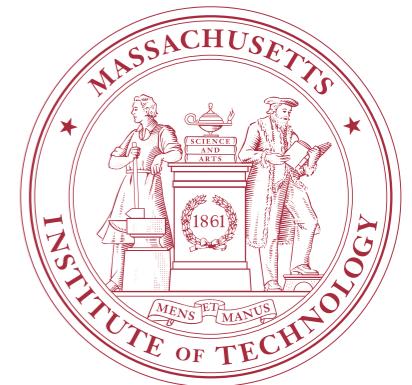
Herman L. Marshall

Norbert S. Schulz, Ralf Heilmann, Sarah Heine (MIT)

David L. Windt (Reflective X-ray Optics)

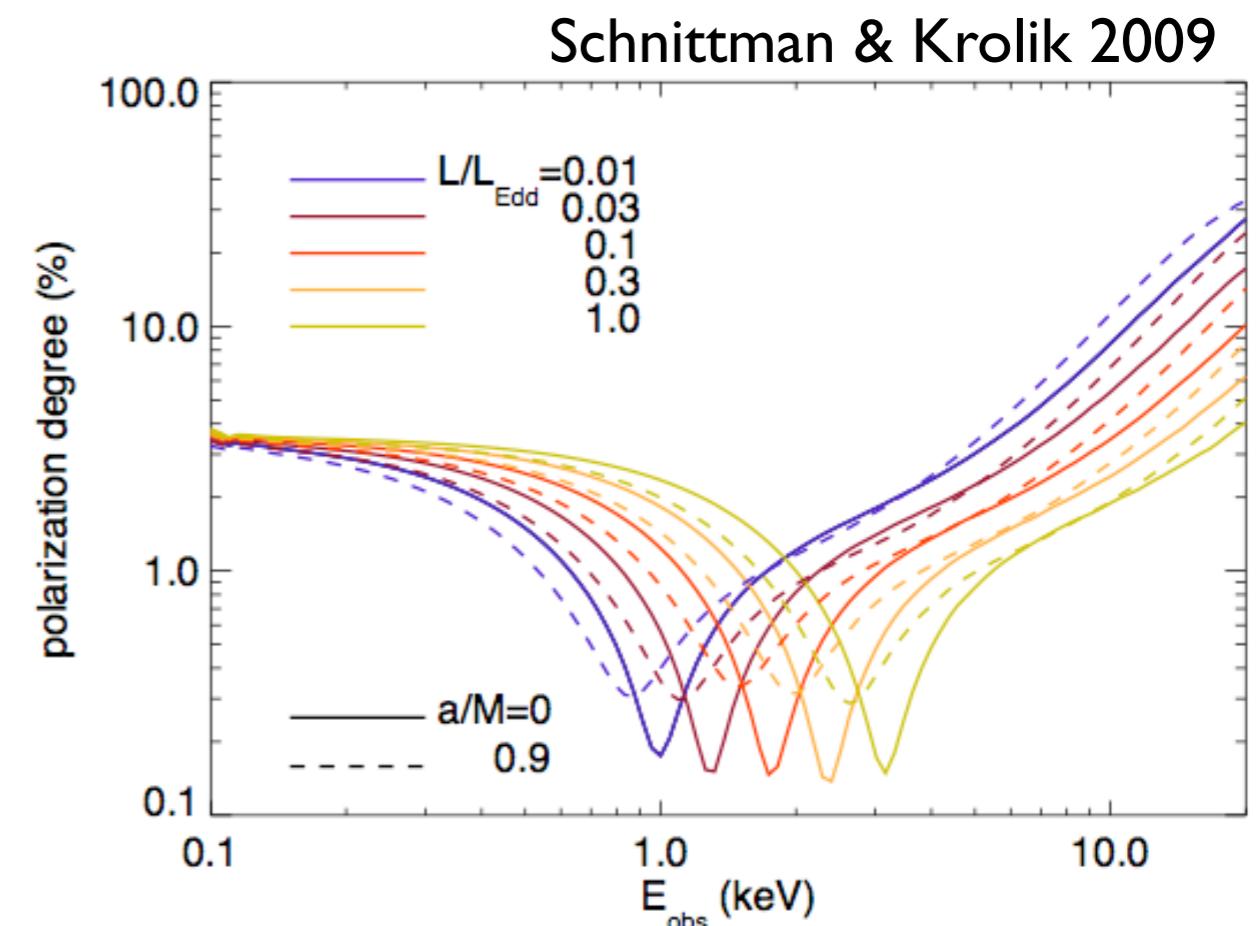
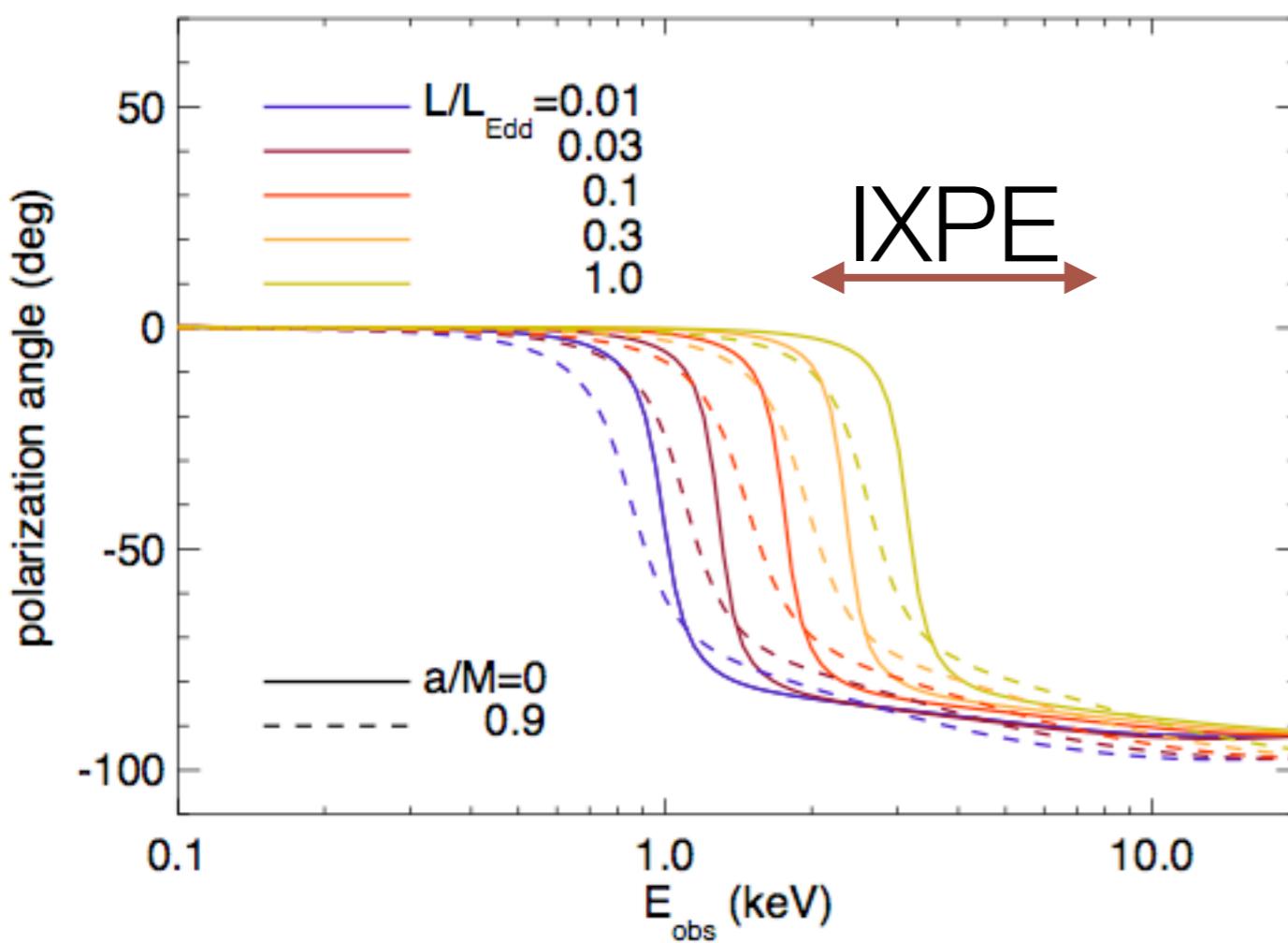
Eric M. Gullikson (LBL)

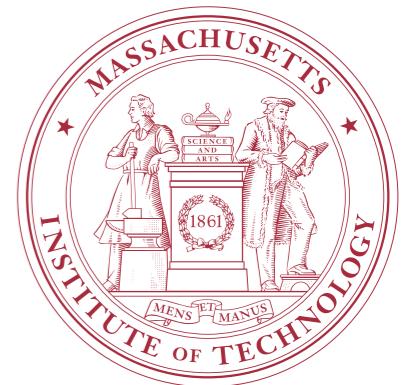
Brian Ramsey (MSFC)



# A Science Goal: Polarimetry of AGN

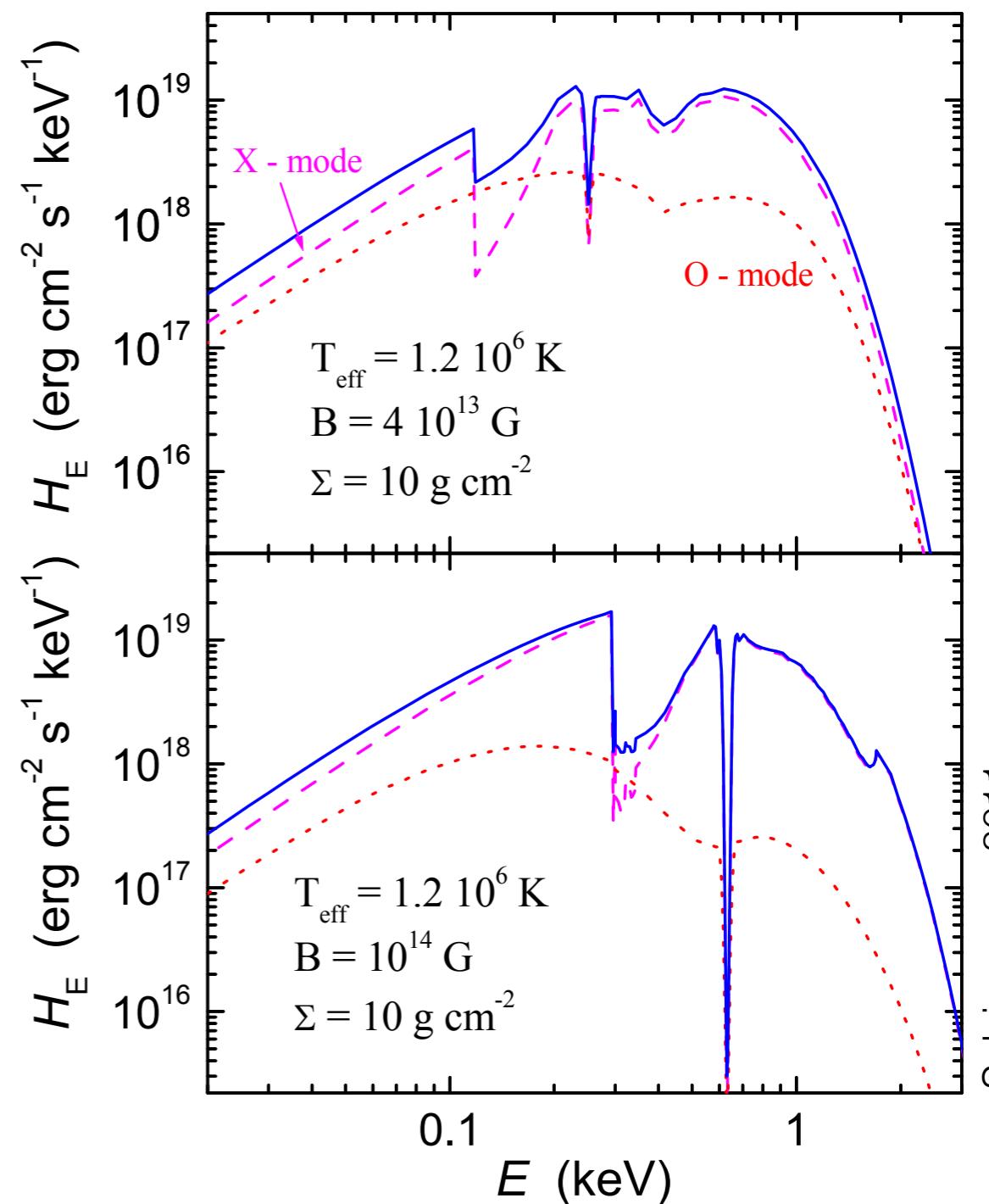
- Scattered return radiation will be polarized
- Polarization fraction and angle depend on  $a/M$ ,  $L/L_{\text{Edd}}$
- Soft and Hard X-ray measurements are needed



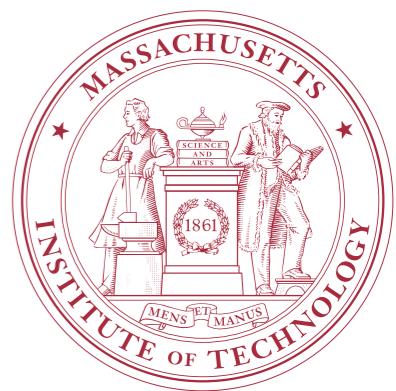


# A Science Goal: Neutron Star Atmospheres

- Isolated neutron stars are often soft, emit below 1 keV
- Polarization amplitude and angle v. rotation phase depends on B-field direction and N-star orientation
- Atmospheres show features now found in spectra of isolated N-stars
- Polarization data would distinguish features in spectra
- Atmosphere models are used to determine  $R^2$ , g to give M,R
- Best target is  $\sim$ 10 mCrab



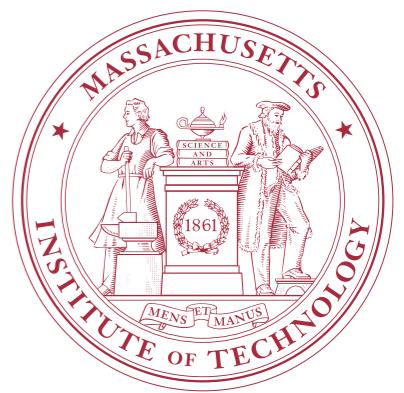
Suleimanov+ 2011



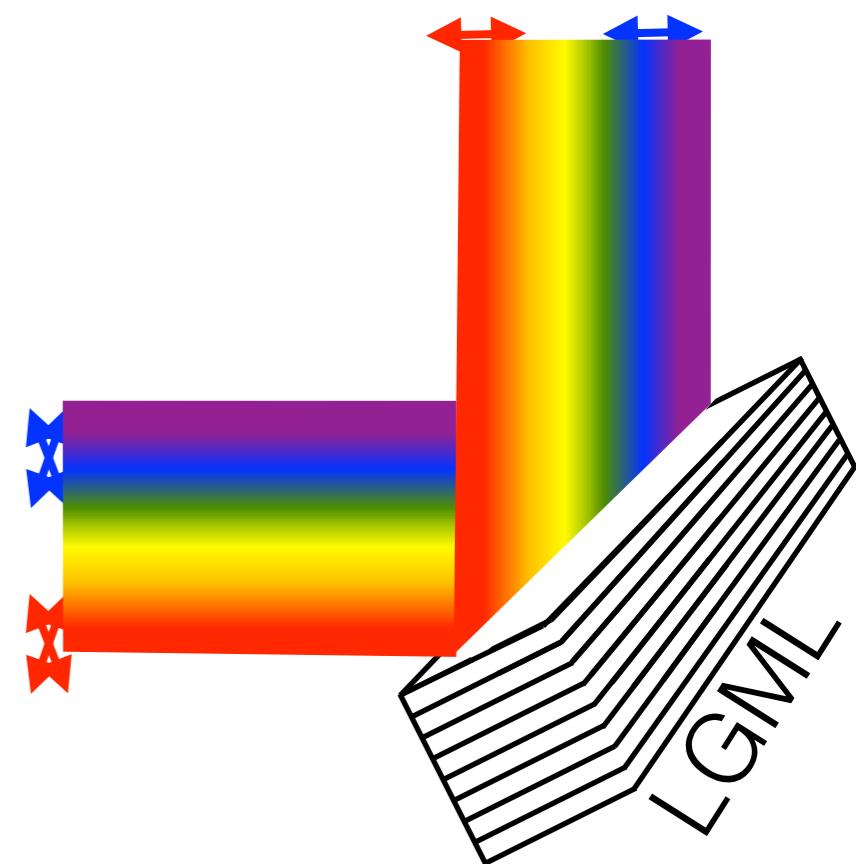
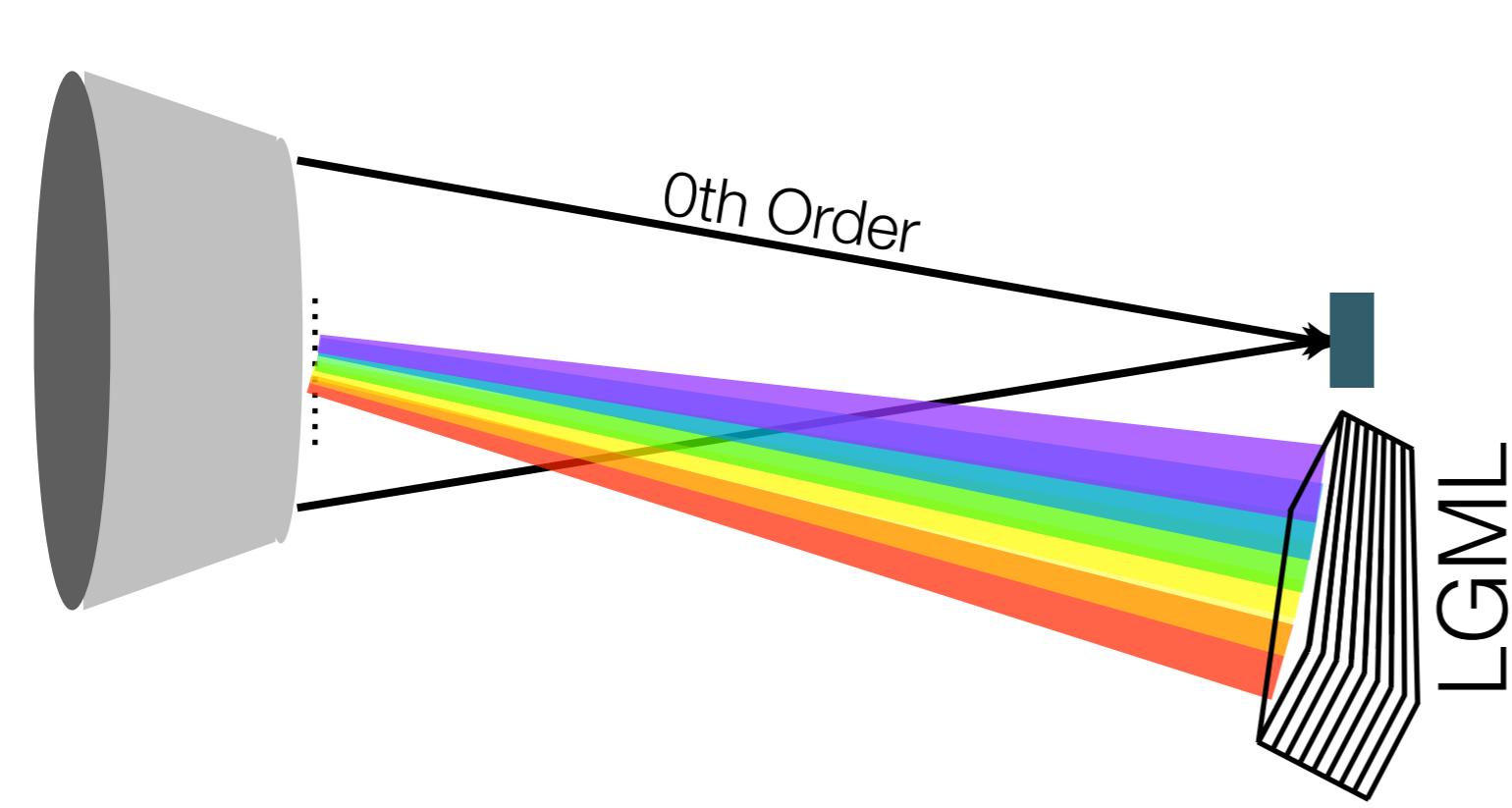
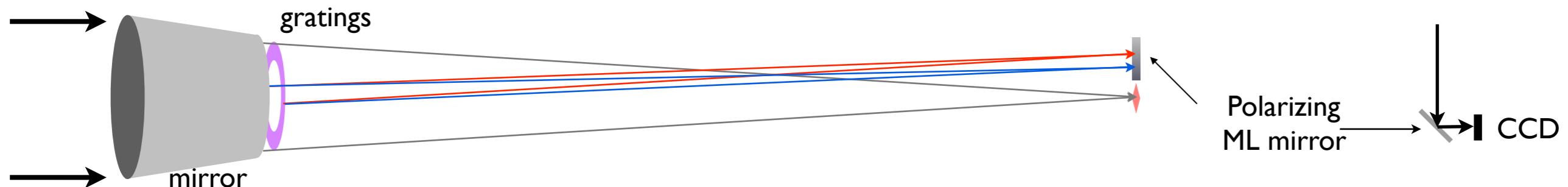
# Soft X-ray Polarimetry

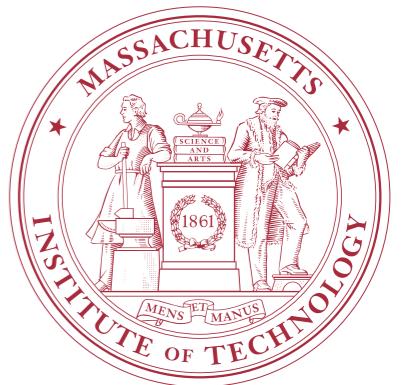
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- Start with general use broad-band focusing mirror
- Use gratings to disperse X-rays: Critical Angle Transmission (CAT) gratings
- Add Bragg reflector: multilayer coated, flat mirror
  - Period is laterally graded (LGML) along mirror to match dispersion
  - Set at  $45\pm5^\circ$  angle (Brewster angle) for >95% selection of s-polarization
- Useful bandpass: 0.15-0.7 keV
- Components verified at MIT Polarimetry Beamline
- Designed a sounding rocket instrument
  - Rocket Experiment Demonstration of a Soft X-ray Polarimeter (REDSoX Polarimeter)
  - Received funding for raytracing and mechanical engineering
  - Resubmitting to NASA APRA

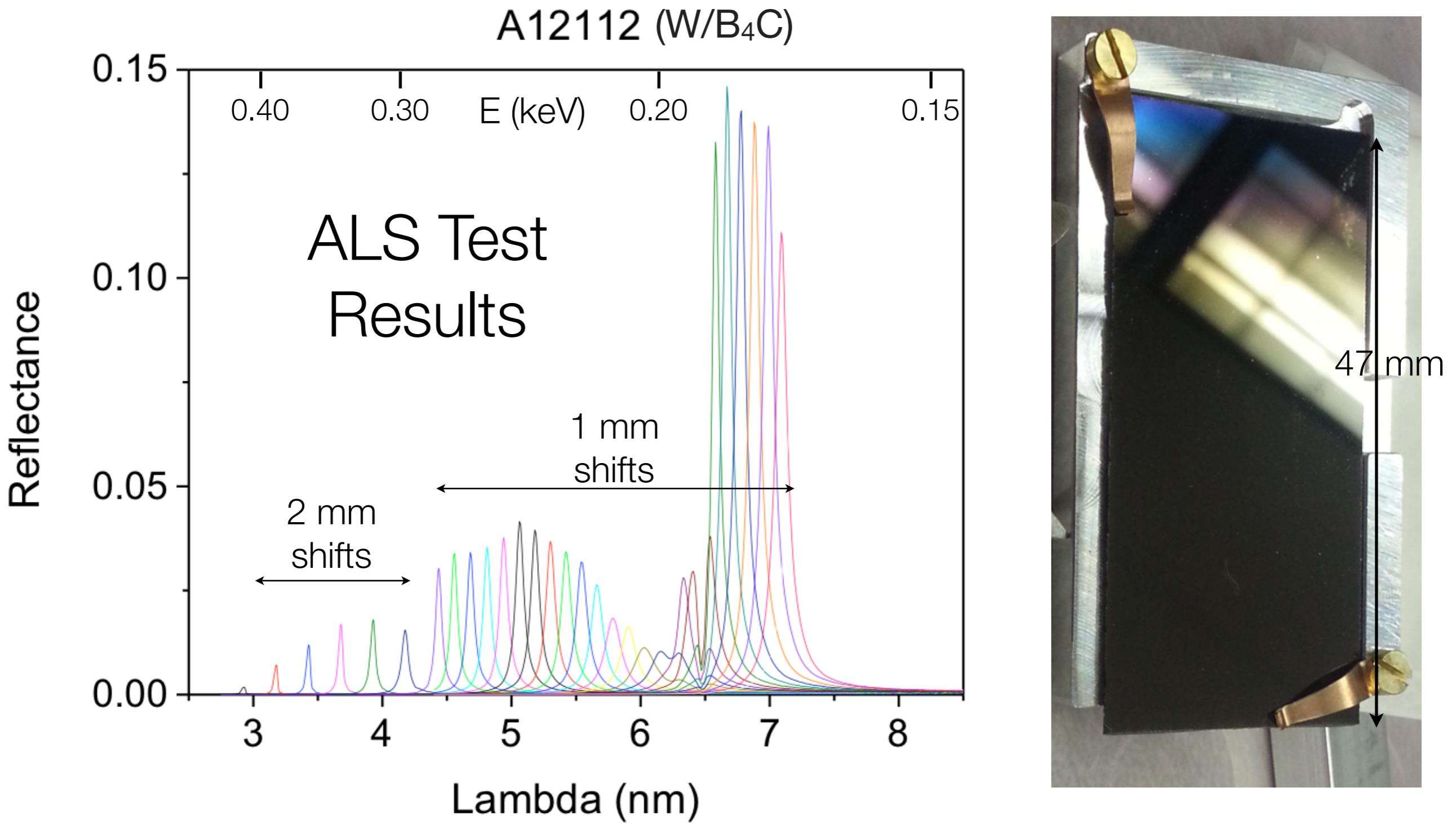


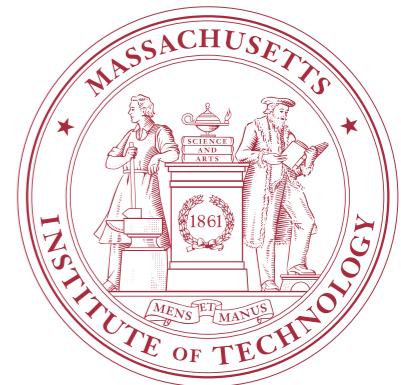
# Multilayer Polarimeter Schematic Layout



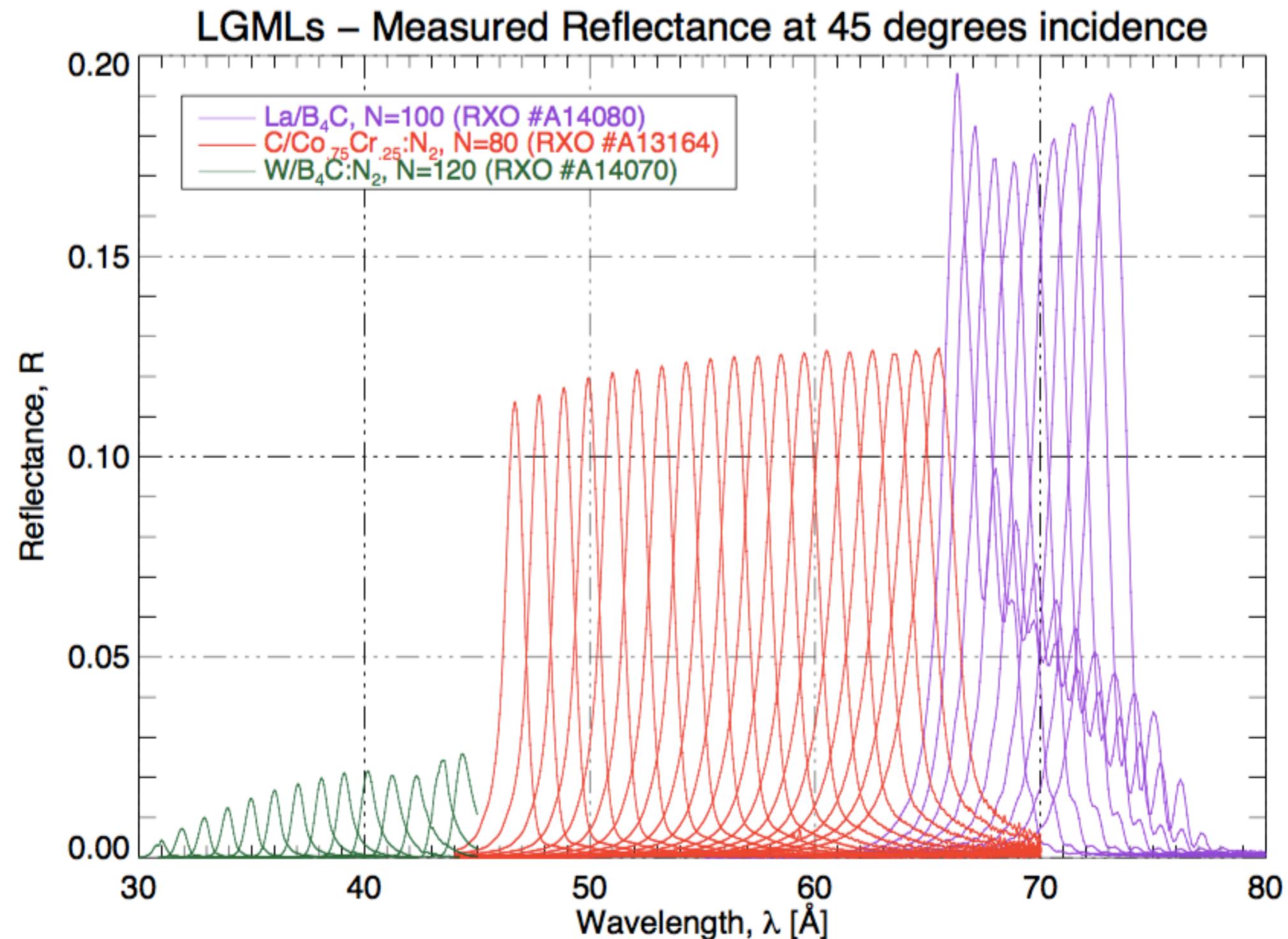


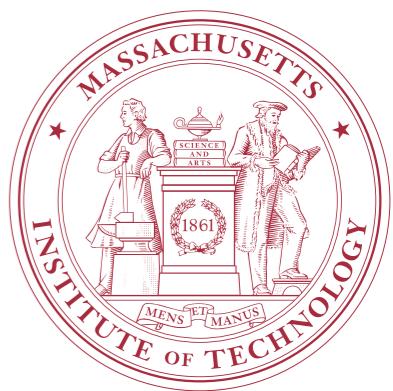
# Laterally Graded Multilayers (LGMLs)





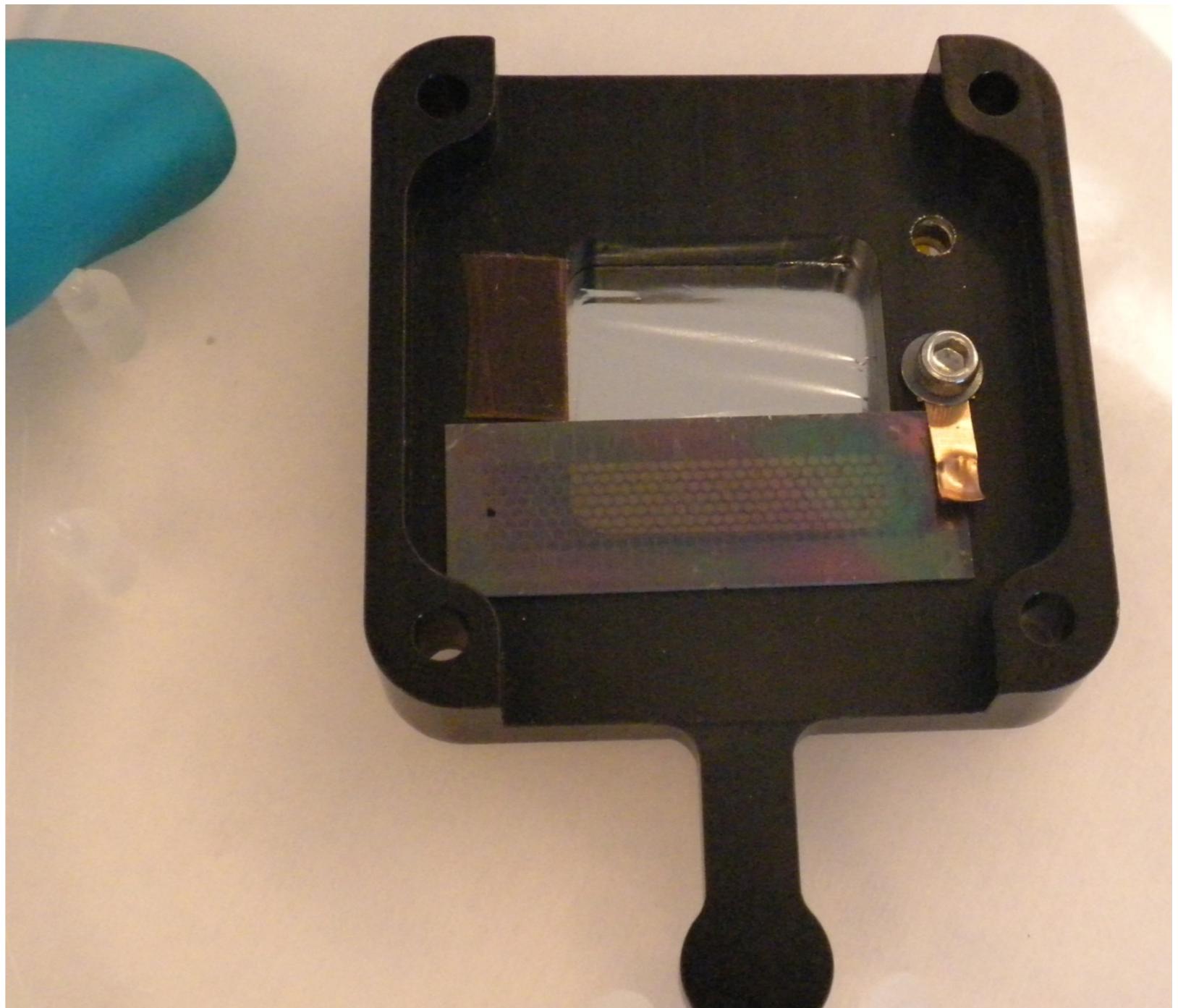
# Current Reflectivities of LGMLs

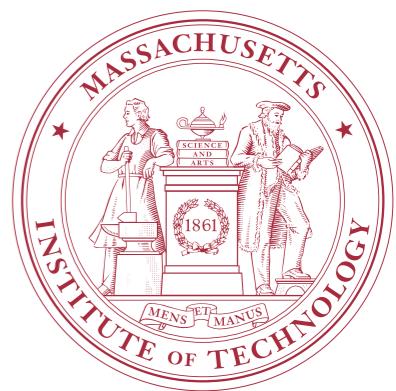




# CAT Gratings

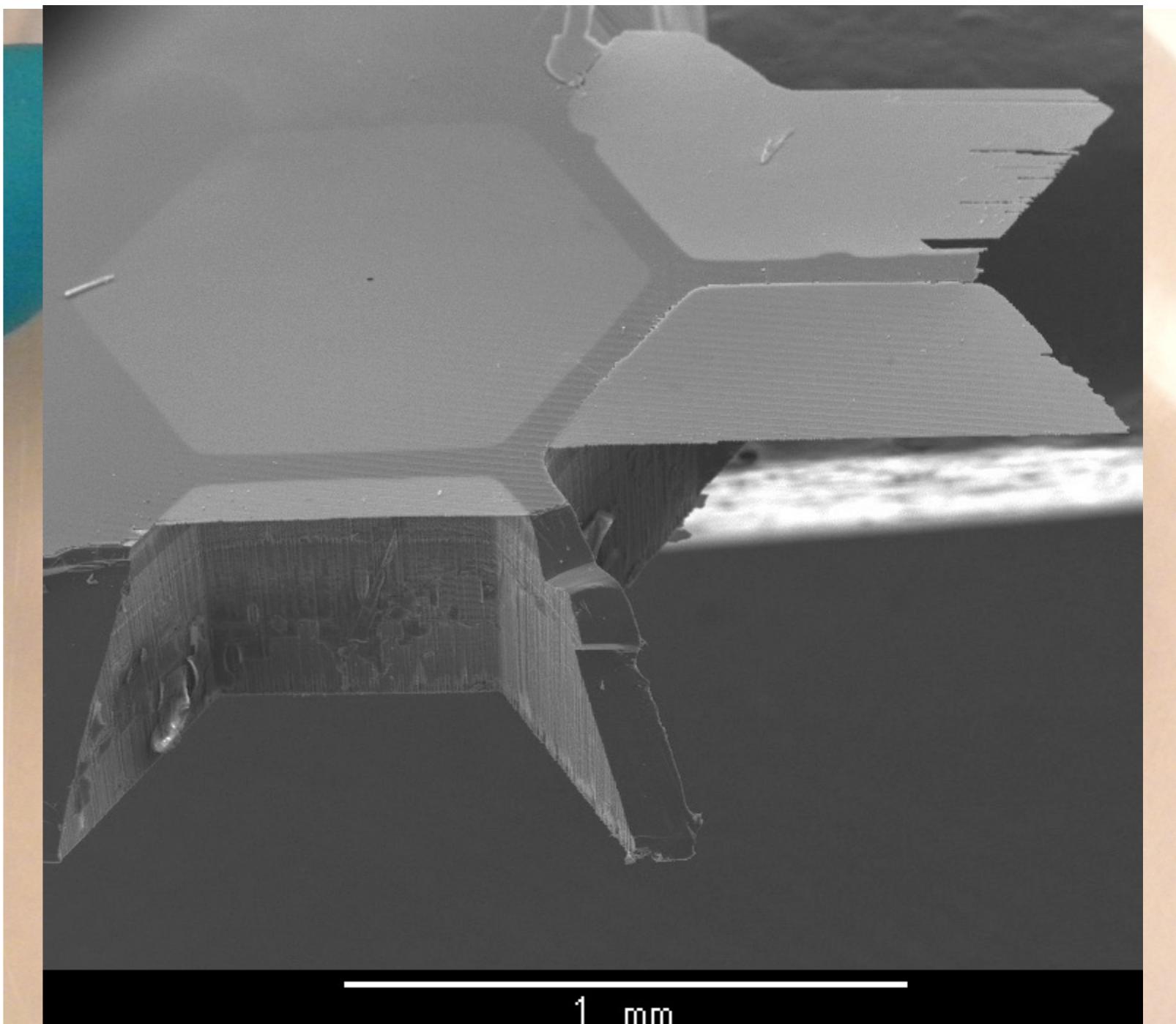
- MIT/SNL CAT gratings are now available
- Gratings etched from Si wafers
- 5 gratings are in lab, others tested at MSFC
- Efficiencies are as expected, up to 25%
- Gratings are light

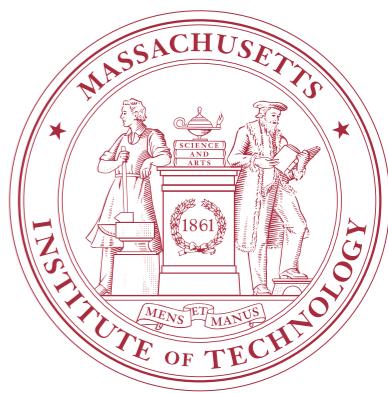




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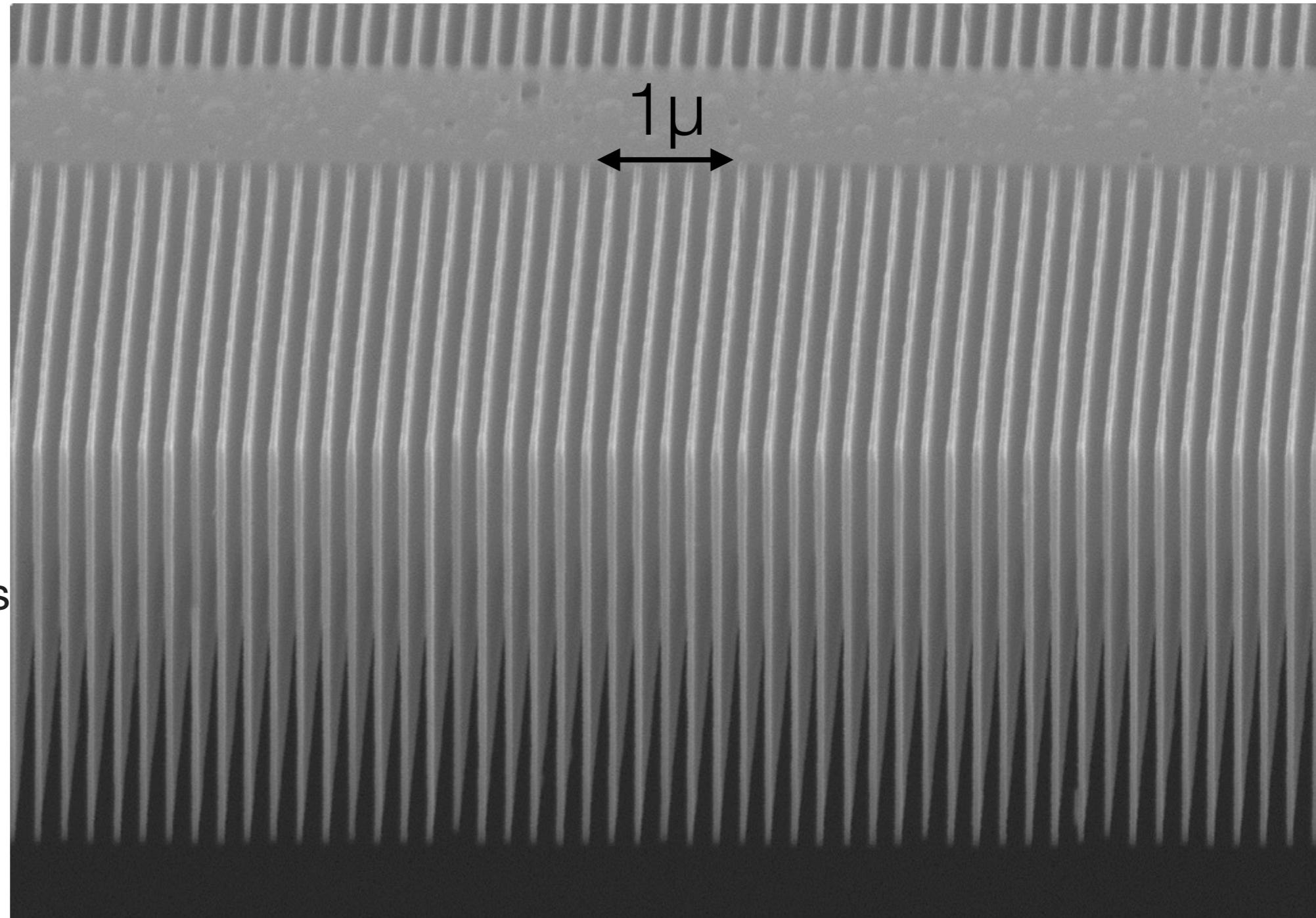
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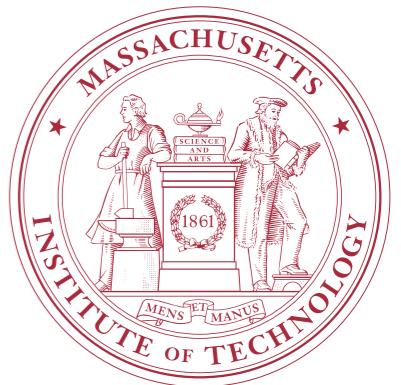




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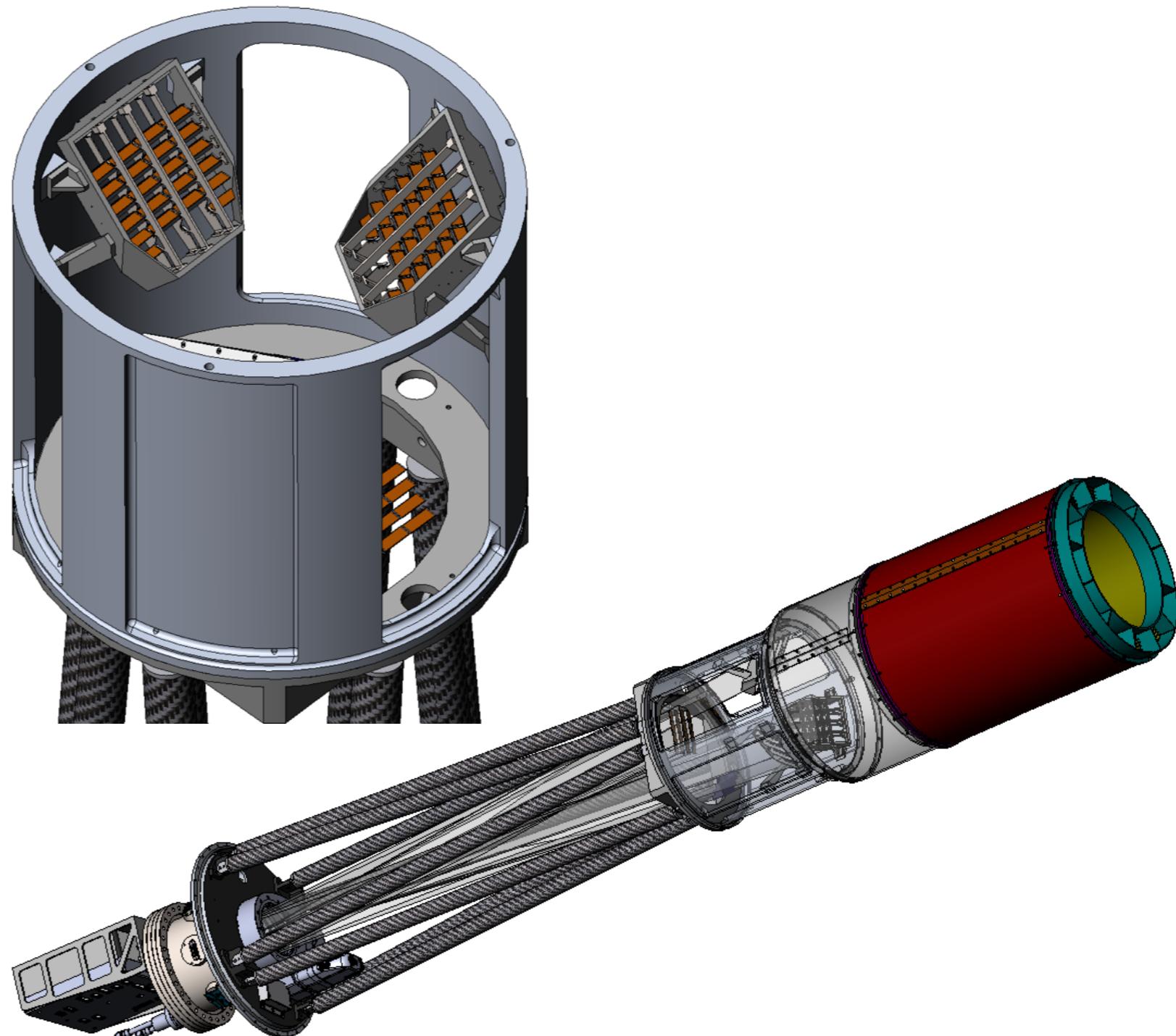
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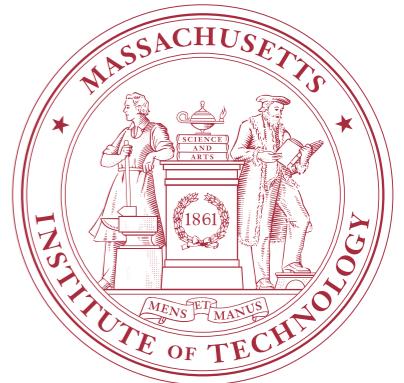




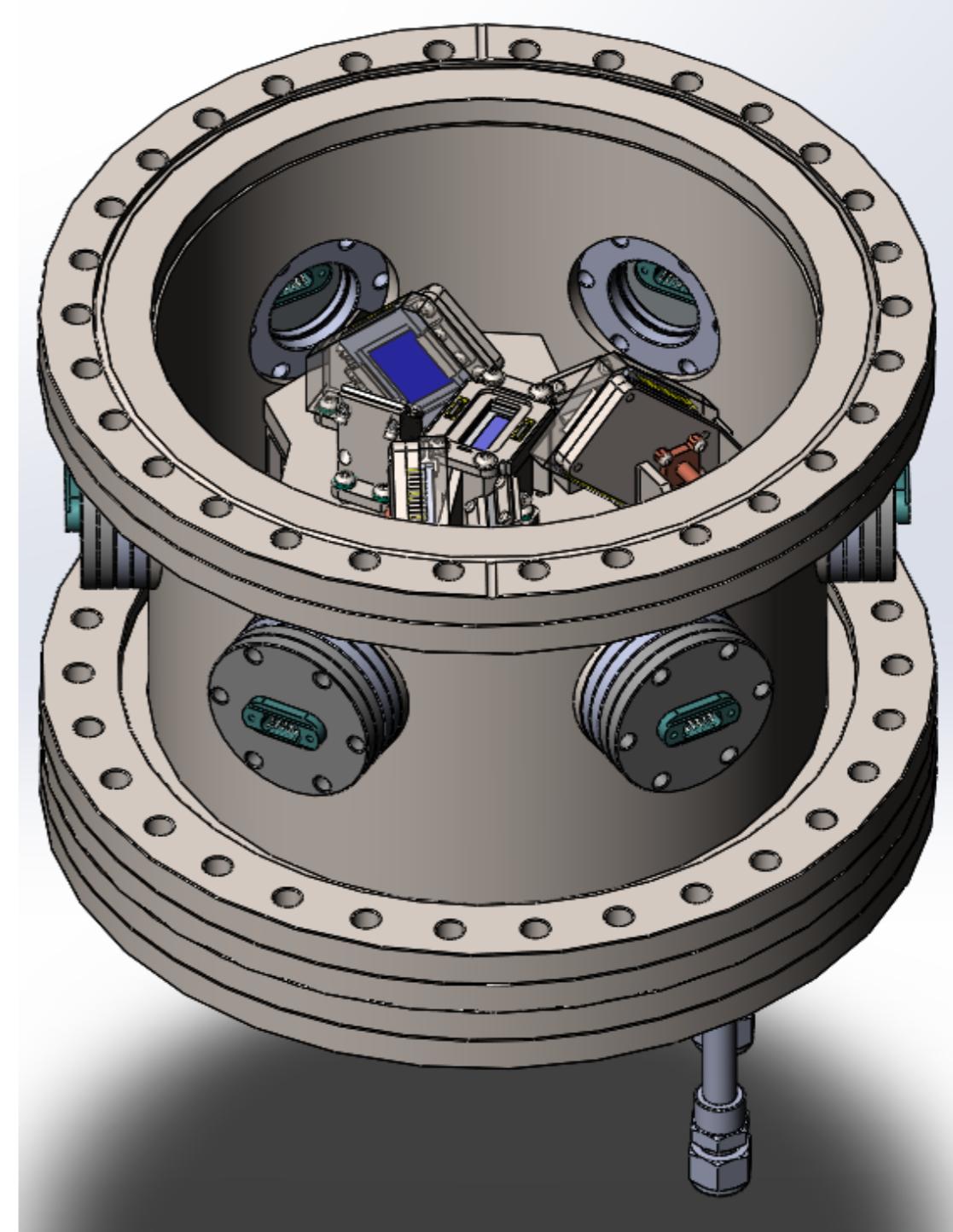
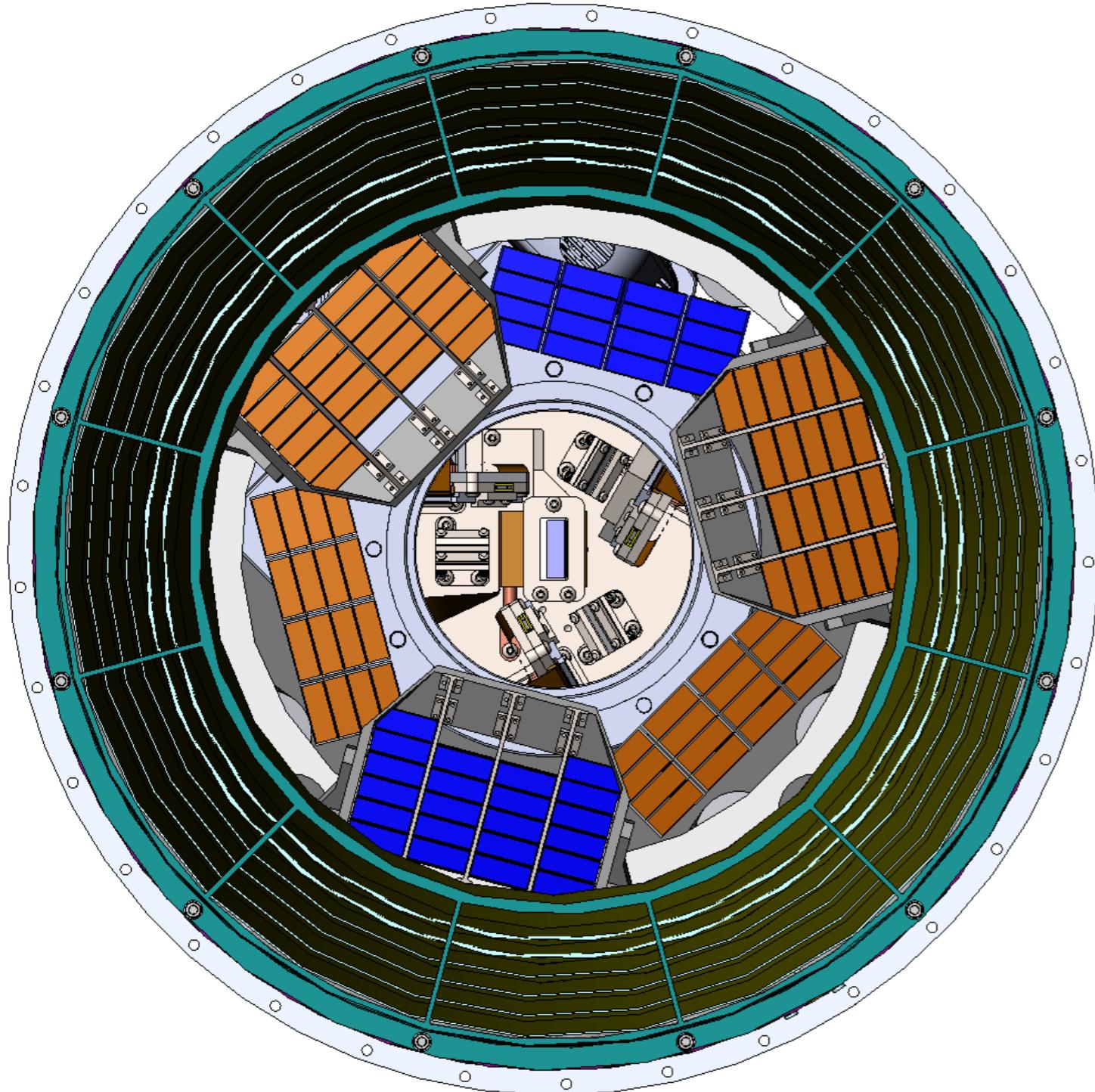
# Sounding Rocket Experiment

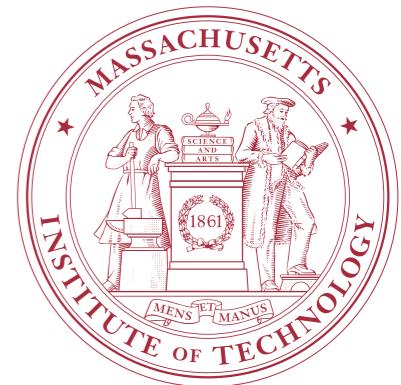
- Focusing optics from MSFC
  - 500 cm<sup>2</sup> mirror
  - 25" HPD
  - $F = 2.5 \text{ m}$ , 44 cm diam.
- CAT gratings (MIT SNL)
  - in 60° sectors
- 1 direct imaging EMCCD (XCAM)
- 3 LGMLs (RXO) & EMCCDs
- Achieve **MDP = 10% for Mk 421** in **300 s exposure**





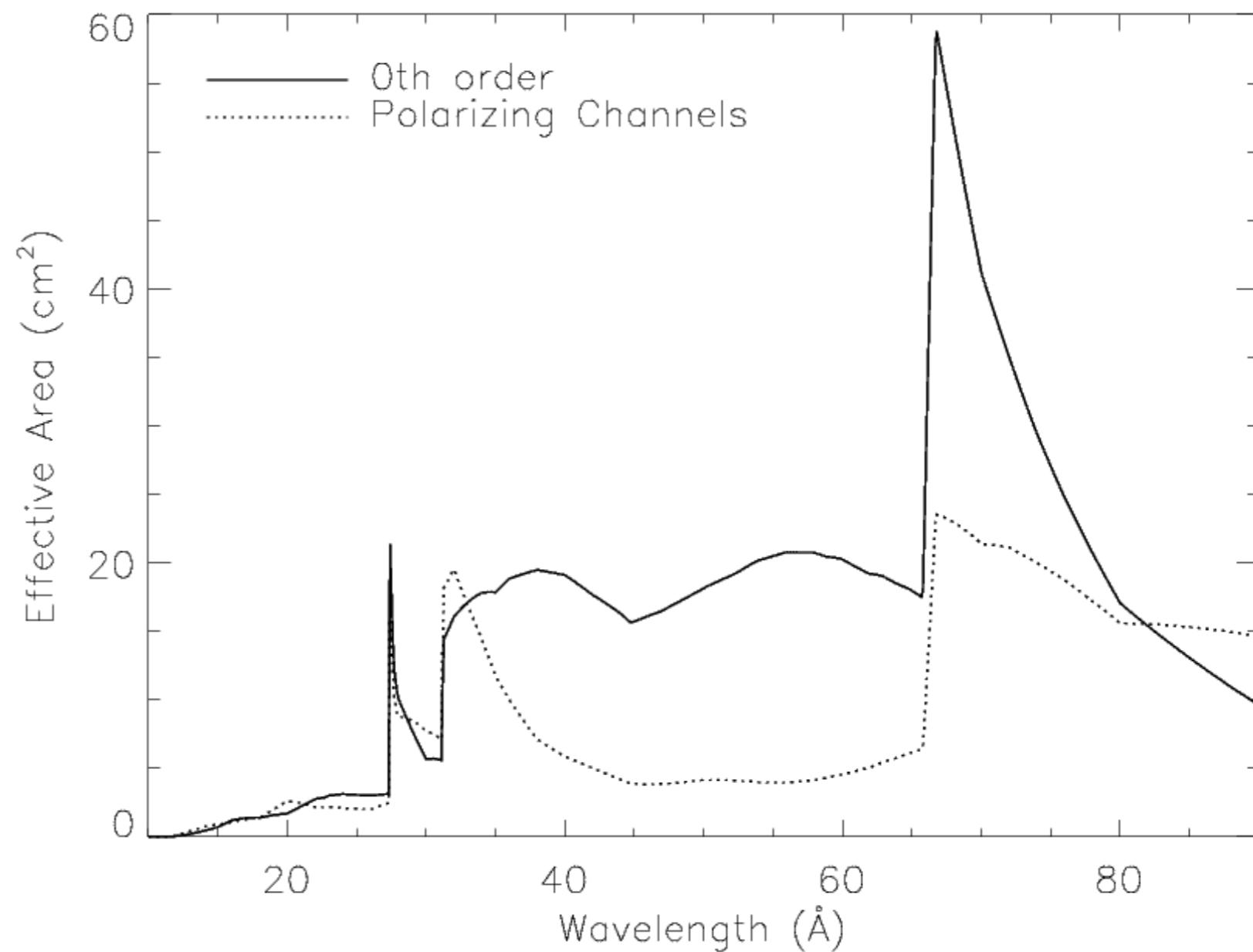
# Suborbital Mission: Entrance Aperture & Focal Plane

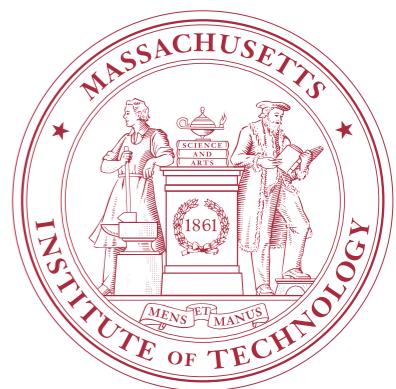




# Prediction for eXTP version

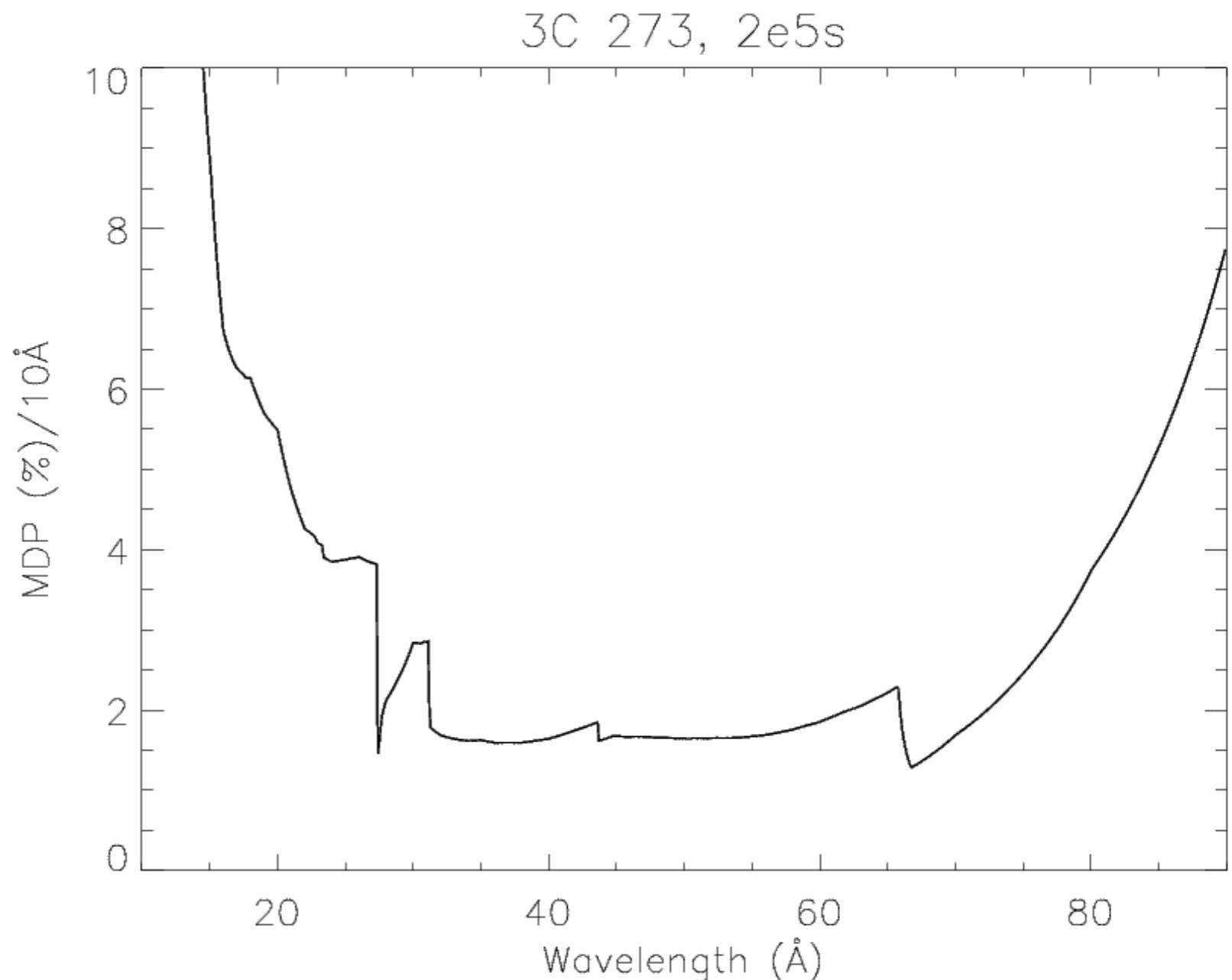
- Same focal plane
- Added improved LGML
- Added mirror shells
- 200 ks exposure on 3C 273
- MDP = 2% in 10Å bands  
(30-70 Å, 0.18 to 0.4)
  - Total rate: 2.3 cps
- PFA in 200 ks: MDP = 2%  
(2-8 keV)

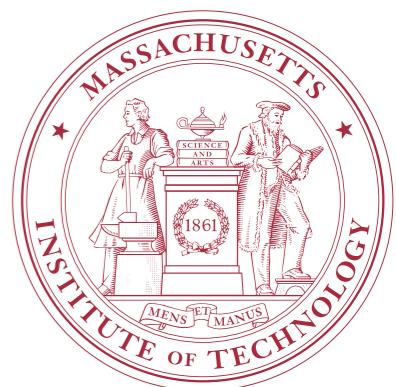




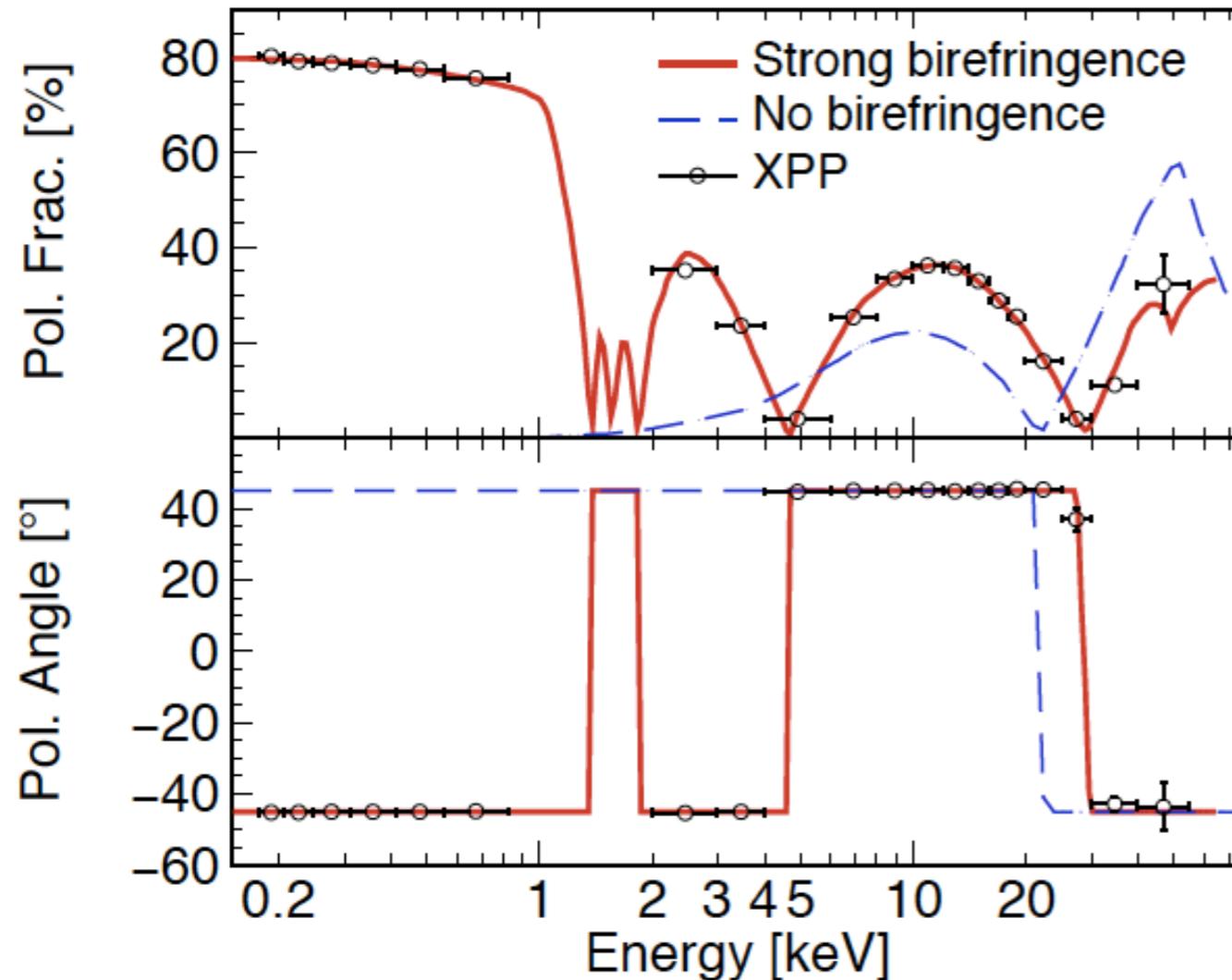
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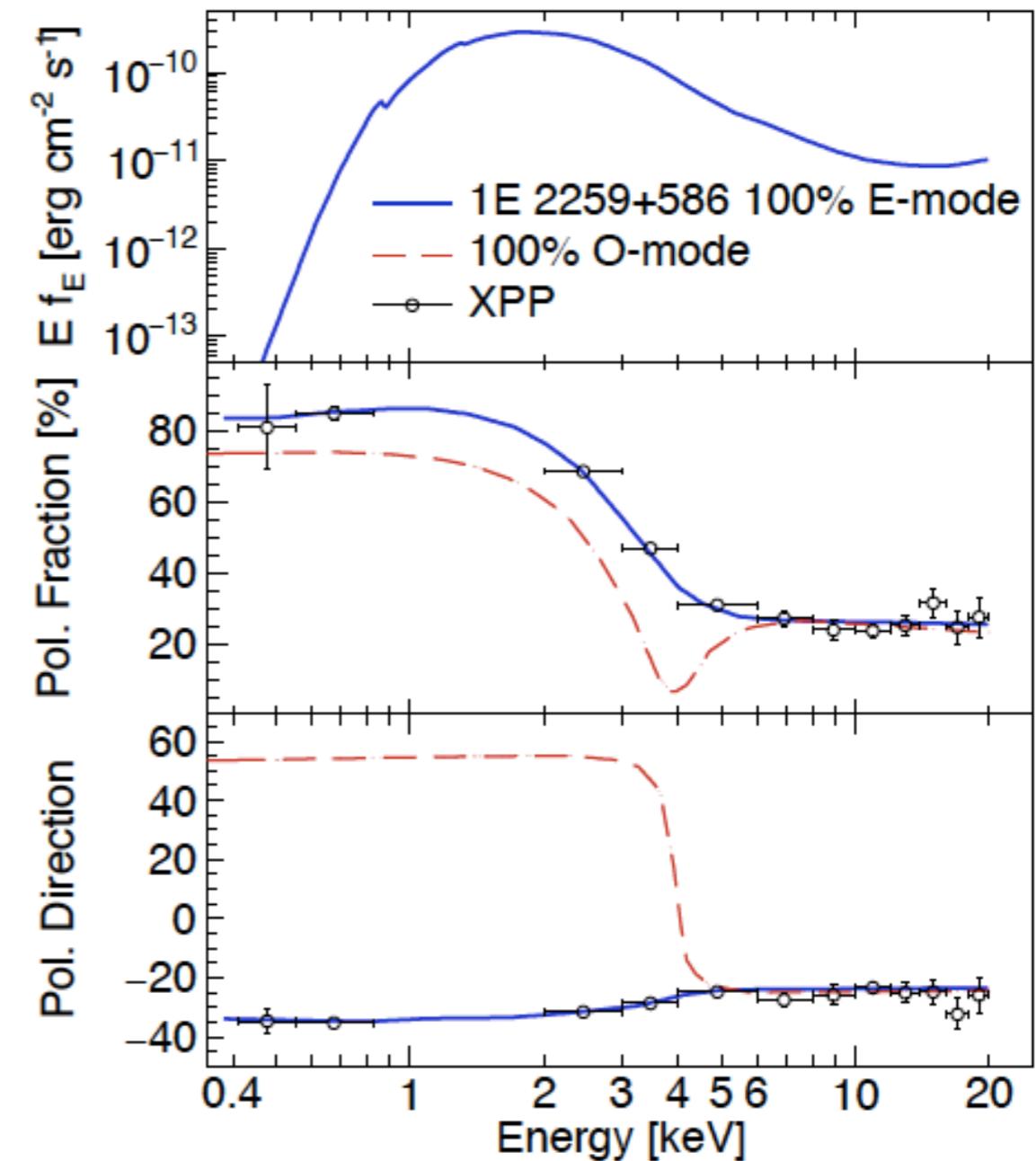




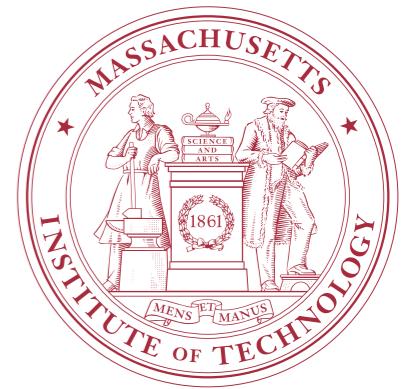
# Polarimetry Probe Science (Krawczynski+ proposal)



Her X-1, 22 ks



1E 2259+586, 1 Ms, 1/4 phase



# Summary

- Soft X-ray (0.15 - 0.7 keV) polarimetry is feasible
  - Suborbital flight proposal: REDSoX Polarimeter (1st launch: 2021)
  - Mech. eng. & raytracing validates approach & achievable tolerances
- NASA probe concept study submitted (Krawczynski+): 0.2-50 keV
- For eXTP: add an instrument of the REDSoX design
  - Replace a SFA? Optics diameter same as SFA and PFA telescopes
  - Add to PFA? Add gratings, 3 LGMLs, & 3 detectors, leave GPD
  - Adds 0.15-0.7 polarimetry to PFA's 2-8 keV band
  - Adds low E imager, spectrometer, and flux monitor
- New/improved eXTP science
  - Blazar jets, quasar cores: polarization over x50 in E
  - Nstars below 1 keV: QED (RX J0720-31), pulsars (Her X-1, PSR 0656+14)
  - AGN disk orientation & GR effects, supporting PFA modeling
  - XRB jets for low NH in BH transients (XTE J1118+480)